

# A ‘constant Lagrangian’ fit of the galaxy rotation curves of the ‘F-series’ from the SPARC database.

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In this paper I apply the ‘constant Lagrangian’ model for galactic dynamics to the F-series of the SPARC database. I will fit the experimental rotation curves of the 16 ‘F’ galaxies from this database using the dual fit approach. This means that one fit is made for the stars dominated region of one galaxy. Another fit is added for the gas dominated region of the same galaxy. The dual fit approach results in a rotation curve fit that mostly remains within the observational error margins. For some galaxies of the series a single model already remains nicely within the error margins and then doesn’t really justify a dual fit. Some galaxies are more complicated and probably need a threefold fit. This dual fit approach follows the mass composition of galaxies as composed of a bulge, possibly a disk and mostly extended gas clouds.

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## I. INTRODUCTION

In a previous paper I introduced the ‘constant Lagrangian’ model for galactic dynamics (de Haas, 2018a). In two sub-sequential papers I made a first qualitative attempt at fitting real rotational velocity curves using the proposed model (de Haas, 2018c,b). In (de Haas, 2018d) paper I gave a more quantitative analysis by including the error bars of the measured velocity. This approach was applied to a subset of 25 galaxies of the [SPARC database](#), including the error margins, as provided by (Lelli et al., 2016). In that subset, non of the F-series galaxies were included. In this paper, the ‘constant Lagrangian’ fit of the F-series is presented. But as with the previous papers, I start by introducing the ‘constant Lagrangian’ model of galactic dynamics. Then I include the dual fit approach that proved to be the best way to model most galaxies. After that, the fit of the F series is presented. The introduction of the model is a copy of the previous two papers and the presentation of the dual fit approach is copied from the previous one, (de Haas, 2018d). The reason to include these two paragraphs again in this paper is to make this paper as more or less stand alone. The original part of this paper, its justification, is the fit of the F-series galaxies from the SPARC database.

## II. THE ‘CONSTANT LAGRANGIAN’ MODEL FOR GALACTIC DYNAMICS

I start by repeating the essentials of this model, which I then apply to the rotation curve of galaxy NGC 1560. The Lagrangian equation of motion reads

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) - \frac{\partial L}{\partial q} = 0. \quad (1)$$

In classical gravitational dynamics I assume circular orbits with  $\dot{q} = v$  and  $q = r$ . The Lagrangian itself is then given by  $L = K - V$ , with  $V$  the Newtonian potential gravitational energy and  $K$  the kinetic energy. One then gets

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) = \frac{dp}{dt} = F. \quad (2)$$

The other part gives

$$\frac{\partial L}{\partial q} = -\frac{dV}{dr}, \quad (3)$$

so one gets Newton's equation of motion in a central field of gravity

$$F_g = -\frac{dV}{dr}. \quad (4)$$

Further analysis of the context results in the identification of the Hamiltonian of the system,  $H = K + V$ , as being a constant of the orbital motion and the virial theorem as describing a relation between  $K$  and  $V$  in one single orbit but also between different orbits,  $2K + V = 0$ .

The previous analysis is problematic relative to the notion of geodetic motion in General Relativity. The problem can best be described in a semi-relativistic approach using the classical Lagrangian equation of motion for geodetic orbits. The most important aspect of geodetic motion in GR is that it requires no force to move on a geodetic. This has important implications for the Lagrangian equation of motion, because  $F = 0$  on a geodetic. One gets

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) = F_g = 0 \quad (5)$$

and as a consequence also

$$\frac{\partial L}{\partial q} = -\frac{dL}{dr} = 0. \quad (6)$$

As a result, one gets the crucial

$$L = K - V = \text{constant} \quad (7)$$

on geodetic orbits.

This result, the Lagrangian of the system as being the constant of the geodetic motion, is used on a daily basis by many of us because it is used by GNSS systems for the relativistic correction of atomic clocks in their satellites. Let's elaborate this a bit further. If we apply the Schwarzschild metric in polar coordinates, we have (Ruggiero et al., 2008)

$$ds^2 = \left( 1 + \frac{2\Phi}{c^2} \right) c^2 dt^2 - \left( 1 + \frac{2\Phi}{c^2} \right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2. \quad (8)$$

In case of a clock on a circular geodesic on the equator of a central non-rotating mass  $M$  we have  $\frac{dr}{dt} = 0$ ,  $\frac{d\theta}{dt} = 0$ ,  $\sin\theta = 1$  and  $\frac{d\phi}{dt} = \omega$ . We thus get, with  $v_{orbit} = r\omega$ , the GR result

$$\frac{d\tau}{dt} = \sqrt{1 + \frac{2\Phi}{c^2} - \frac{v_{orbit}^2}{c^2}} \quad (9)$$

with  $d\tau$  as the clock-rate of a standard clock A in a geodetic orbit and  $dt$  as the 'universal' clock-rate G of a standard clock at rest in infinity, the only condition for which  $d\tau = dt$ .

The result of Eqn. (9) is the basic relativistic correction used in GNSS clock frequencies, with the first as the gravity effect or gravitational potential correction and the second as the velocity effect or the correction due to Special Relativity (Ashby, 2002; Hećimović, 2013; Delva and Lodewyck, 2013).

Given the classical definitions of  $K = \frac{mv_{orbit}^2}{2}$  and  $V = m\Phi$ , we get

$$\frac{d\tau}{dt} = \sqrt{1 - \frac{2L}{U_0}}. \quad (10)$$

All the satellites of a GNSS system are being installed on a similar orbit and thus syntonized relative to one another because they share the same high and velocity and have constant  $L$  and  $\frac{d\tau}{dt}$  on those orbits. But different GNSS systems, as for example GPS compared to GALILEO, are functioning on different orbits with different velocities and those systems aren't syntonized relative to one another. This non-syntonization between satellites on orbits with different heights and virial theorem connected velocities is very annoying for the effort towards realizing an integration of the different GNSS systems into one single global network.

Fundamental in the approach of this paper is to analyze gravity using relative frequency shifts, and thus  $\frac{d\tau}{dt}$ , as one of the basic experimental inputs. Such a method is also looming in today's geodesy. In modern gravitational geodesy scientists are investigating the relativistic frequency shift as a new observable type for gravity field recovery (Mayrhofer and Pail, 2012). Driven by this development, modern geodesy is about to go through a change from the Newtonian paradigm to Einstein's theory of general relativity (Kopeikin et al., 2017). A new generation of atomic clock is the game changer for this new domain of chronometric geodesy, and requires additional new techniques to be developed in the field of frequency transfer and comparison (Delva and Lodewyck, 2013). The paradigm shift towards gravitational divergence recovery is based on the principle of frequency comparison between two clocks on different space-time locations in order to measure the frequency shift between them (Delva and Lodewyck, 2013). The knowledge of the Earth's gravitational field has often been used to predict frequency shifts between distant clocks. In relativistic geodesy, the problem is reversed and the measurement of frequency shifts between distant clocks now provides knowledge of the gravitational field (Delva and Lodewyck, 2013). This reversal also looms in my postulate of the 'constant Lagrangian' model. A constant Lagrangian implies a zero divergence in the syntonization of atomic oscillators and thus an absence of gravitational

stress. A divergence in the Lagrangian implies a divergence in the time dilation factor  $\frac{dt}{dr}$  and thus a non-zero gravitational stress.

The ‘constant Lagrangian’ model for galactic dynamics starts with the postulate that the geodetic Lagrangian  $L = K - V$  is a galactic constant, not just an orbital constant. In this ‘time bubble halo’ model the classical Newtonian potential is assumed valid. This potential in the case of a model galaxy with a perfect quasi-solid bulge and a perfect Schwarzschild emptiness around it is given in Fig.(1). My model galaxy is build of a model bulge with mass  $M$  and radius  $R$  and a Schwarzschild metric emptiness around it. The model bulge has constant density  $\rho_0 = \frac{M}{V} = \frac{3M}{4\pi R^3}$  and its composing stars rotate on geodetics in a quasi-solid way. So all those stars in the bulge have equal angular velocity on their geodetic orbits, with  $v = \omega r$ . On the boundary between the quasi solid spherical bulge and the emptiness outside of it, the orbital velocities are behaving smoothly. So the last star in the bulge and the first star in the Schwarzschild region have equal velocities and potentials. I also assume that the Newtonian potential itself is unchanged and unchallenged, remains classical in the whole galaxy and its surroundings. Such a model galaxy doesn’t have a SMBH in the center of its bulge and it only has some very lonely stars in the space outside the bulge.

The ‘constant Lagrangian model postulates  $L = K - V = constant$  in the entire galaxy, without changing the Newtonian potential. As a result, in such a model bulge,  $L$  is a constant of the motion, not only in one orbit but also between orbits.

$$\frac{L}{m} = \frac{v_{orbit}^2}{2} + \frac{GM}{r} = \frac{3GM}{2R} = constant. \quad (11)$$

For the region  $0 \leq r \leq R$  we get

$$v_{orbit}^2 = \frac{GM}{R} \cdot \frac{r^2}{R^2} \quad (12)$$

and outside the model bulge, where  $R \leq r \leq \infty$ , we have

$$v_{orbit}^2 = \frac{3GM}{R} - \frac{GM}{r}. \quad (13)$$

From the perspective of a free fall Einstein elevator observer, the free fall on a radial geodetic from infinity towards the center of the bulge, the other free fall tangential geodetics seem to abide the law of conservation of energy, because the escape kinetic energy plus the orbital kinetic energy is a constant on my model galaxy with galactic constant  $L$ . An Einstein elevator system with test mass  $m$  that would be put in an orbital collapse situation, magically descending from orbit to orbit in a process in thermodynamic equilibrium, would

Point	Relation	Expression
Outside the <b>bulge</b>	$r > R$	$-\frac{GM}{r}$
On the Surface	$r = R$	$-\frac{GM}{R}$
Inside the <b>bulge</b>	$r < R$	$-GM \left[ \frac{3R^2 - r^2}{2R^3} \right]$
At the centre	$r = 0$	$-\frac{3}{2} \left( \frac{GM}{R} \right)$

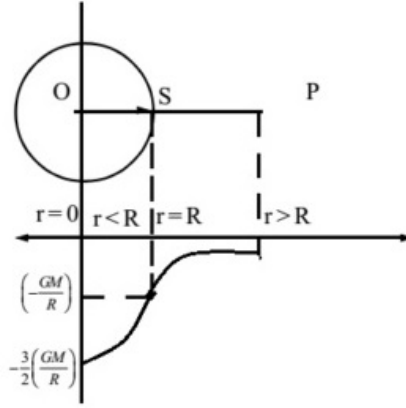


FIG. 1. The potential inside and out of a model bulge

have constant total kinetic energy, from the radial free fall perspective. This can be expressed as  $L = K_{orbit} - V = K_{orbit} + K_{escape} = K_{final}$ . In Fig.(2) I sketched the result, with  $-V = +K_{escape}$ .

Such a model galaxy would also be a GNSS engineer's dream come true because the whole model galaxy is in one single syntonized time-bubble.

$$\frac{d\tau}{dt} = \sqrt{1 - \frac{2L}{U_0}}. \quad (14)$$

Given the Baryonic Tully-Fisher relation in Milgrom's version  $v_{final}^4 = Ga_0M$  with  $2\pi a_0 \approx cH_0$ , with  $a_0$  as Milgrom's galactic minimum acceleration and  $H_0$  as the Hubble constant (Milgrom, 1983; McGaugh, 2005), we get as a galactic time bubble fix

$$\frac{d\tau}{dt} = \sqrt{1 - \frac{2L}{U_0}} = \sqrt{1 - \frac{v_{final}^2}{c^2}} = \sqrt{1 - \sqrt{\frac{v_{final}^4}{c^4}}} = \quad (15)$$

$$\sqrt{1 - \sqrt{\frac{Ga_0M}{c^4}}} = \sqrt{1 - \sqrt{\frac{GH_0M}{2\pi c^3}}} = \sqrt{1 - \sqrt{\frac{M}{2\pi M_U}}}, \quad (16)$$

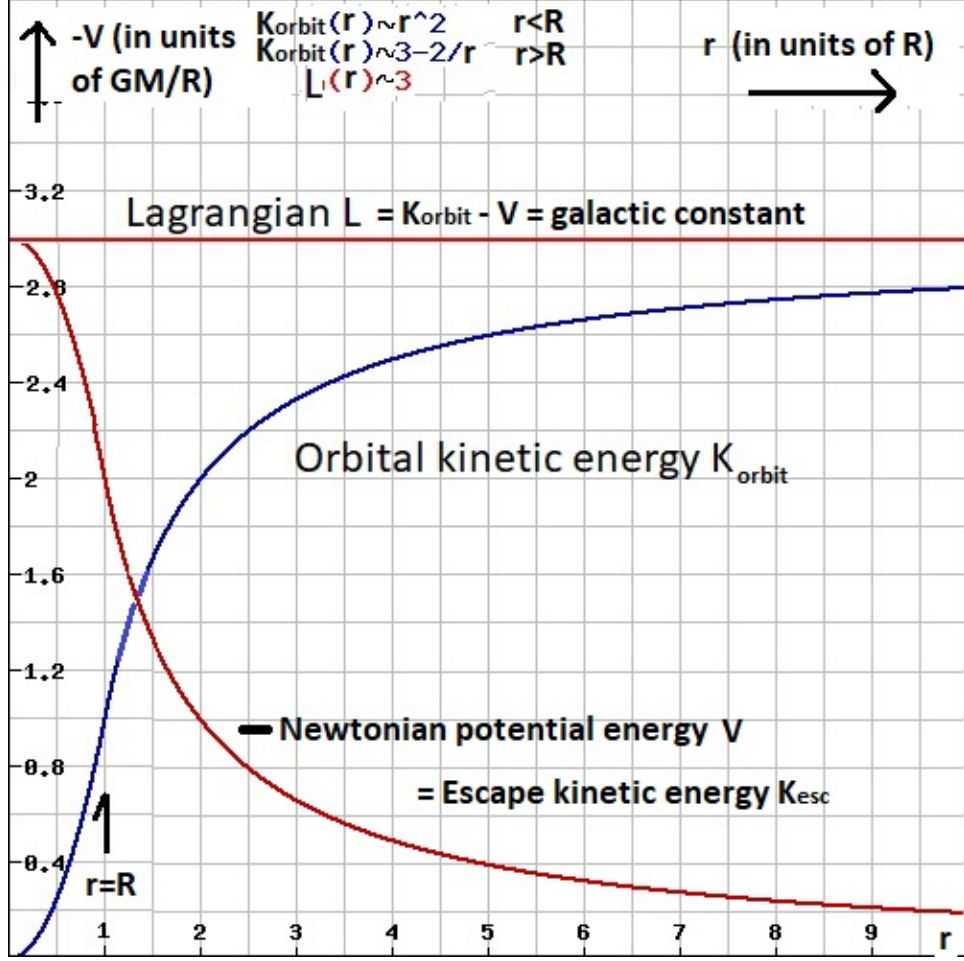


FIG. 2. The square of the orbital velocity profile in the model galaxy with  $L = constant$ .

in which I used  $L = 3GM/R = K_{final} = \frac{1}{2}mv_{final}^2$  and  $M_U = \frac{c^3}{GH_0}$ . This last constant can be referred to as an apparent mass of the Universe, a purely theoretical number constant, see (Mercier, 2015).

In a model Universe, this would imply that my model galaxy would be in a proper time bubble with clock-rate  $d\tau$  relative to the universal clock-rate  $dt$  in proportion to the masses of galaxy  $M$  and Universe  $M_U$ . In my model galaxy theoretical environment the Baryonic Tully-Fisher relationship implies that the galactic time bubble is fixed through the mass of my model galaxy and that this fix is a cosmological one. So what is a universal acceleration minimum  $a_0$  in MOND can be interpreted as a universally correlated (through  $M_U$ ) but still local (through  $M$ ) time bubble fix in my model galaxy geodetic environment. In such a model Universe, the time bubble of a galaxy immediately functions as a gravitational lens, because  $\frac{d\tau}{dt}$ , as measuring the curvature of the metric, also determines the gravitational index



of refraction of the time bubble relative to Cosmic space where  $dt = d\tau$ . In my model, the Dark Matter halo is above all present as a time bubble halo, determined by the factor  $\frac{d\tau}{dt}$ .

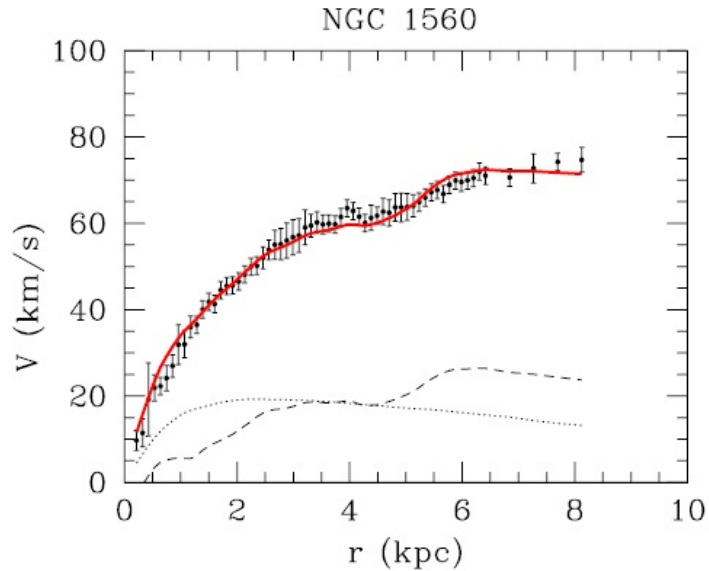
### III. THE DUAL FIT APPROACH TOWARDS ROTATION CURVES

Having determined the model galactic velocity rotation curve based on the Lagrangian as a galactic constant of orbital motion, the question is to what extent real galaxies can be modeled in this way. In my Lagrangian approach I analyze the plot of  $v_{orb}^2$ , in  $(km/s)^2$  against  $r$ , in  $kpc$ . This is in contrast to the usual rotation curves where  $v_{orb}$ , in  $(km/s)^2$  against  $r$ , in  $kpc$ . In the Lagrangian approach, the energies, not the velocities, are primary. In each plot the experimental values are in red stars with vertical error bars and the theoretical values are in black circles and black triangles. The fitting plot is with one single fit for  $M$ , in units of  $10^{10} M_{solar}$ , and  $R$ , in units of  $kpc$ . The most important cut in the model is the change from the model bulge to the model empty space around it. In the model bulge,  $V_{orb}^2 \propto r^2$ , outside the model bulge  $V_{orb}^2 \propto -r^{-1}$ . In the dual approach, another important part is the shift from the first  $(M_1, R_1)$ -fit to the second  $(M_2, R_2)$ -fit.

This shift from a first fit to a second fit for the rotation curve of one single galaxy grew out of the analysis of the data. The justification of this dual approach is situated in the difference between the pure model with a perfect bulge and an emptiness around it with only some spare stars and some H1-gas. In reality, galaxies are way more complicated, as being composed of an inner bulge, a disk and H1 gas clouds. The constant Lagrangian' model includes the classical Newtonian potential and that allows for the use of inner and outer shells in the model approach. A shell of H1-gas that is treated as a perfect outside shell can be ignored for the inner shell dynamics. A shell of H1-gas that is an inner shell can be treated as part of the bulge. Going through a shell means leaving the inner model curve and progressing to the outer model curve.

This experimentally driven extension of the pure model with a perfect bulge and a perfect Schwarzschild empty metric around it appeared first while analyzing the rotation curve of NGC 1560. The velocity rotation curve data of NGC 1560 come from (Broeils, 1992). The comparison with other fitting models came from (Gentile et al., 2010). The rotation curve of NGC 1560 has a peculiar "wobble". This "wobble" in the graph of NGC is described in the following quote.

In the rotation curve of NGC 1560, as derived by B92, there is a clear “wiggle” in the total rotation velocity, which corresponds very closely to a similar wiggle in the gas contribution to the rotation curve. Mass models such as MOND naturally reproduce the feature, whereas models that include a dominant spherical (or triaxial) halo are too smooth to do so. (Gentile et al., 2010)



**Figure 12.** Rotation curve fit using MOND. The best-fit distance is 2.94 Mpc. Lines and symbols are like those in Fig. 10.

FIG. 3. (Gentile et al., 2010) fit of NGC 1560, using MOND. In this graph, the “wiggle” occurs where the stars, represented by the dotted line, give in to the H1 gas clouds, represented by the barred line, around 4 – 5 kpc.

In my approach, I first have to determine the model galaxy curve that fits best, using the parameters  $M$  and  $R$ , and then I can use  $M$  as a free parameter in order to create a perfect time bubble. In case of NGC 1560 however, it seems that in phase 1 two models partially fit the rotation curve. The first pure model fits NGC 1560 before the “wobble” the second pure model fits NGC 1560 after the “wobble”, see Fig.(4). Thus in my approach, the ‘constant Lagrangian’ model, the modeling indicates the underlying dynamics and the real world of unpredictable but observable mass distributions of stars and H1 gas clouds. The “wobble” divides NGC 1560 in two regions, which both follow their respective pure model relatively smoothly without being disturbed by that other part of the galaxy. The baryonic matter further away from the center than the H1 gas of producing “wobble” just behaves as if the bulge end where the “wobble” ends. The baryonic matter closer to the center than the H1 gas of producing “wobble” just behaves as if the H1 gas of the “wobble” is a perfect shell which doesn’t gravitate inside that shell. So the fact that two pure models can be made to partially fit the rotation curve actually reveals a lot of the underlying dynamics, reproducing known baryonic behavior under the influence of a Newtonian potential.

This fact, that one can only reasonably fit NGC 1560 by using two model fits, appeared to be the case for most galaxies. The shift from one model fit to the other model fit mostly follows the shift from stars dominated mass distribution to H1 gas dominated mass distribution. Sometimes however it follows the shift from bulge to disk.

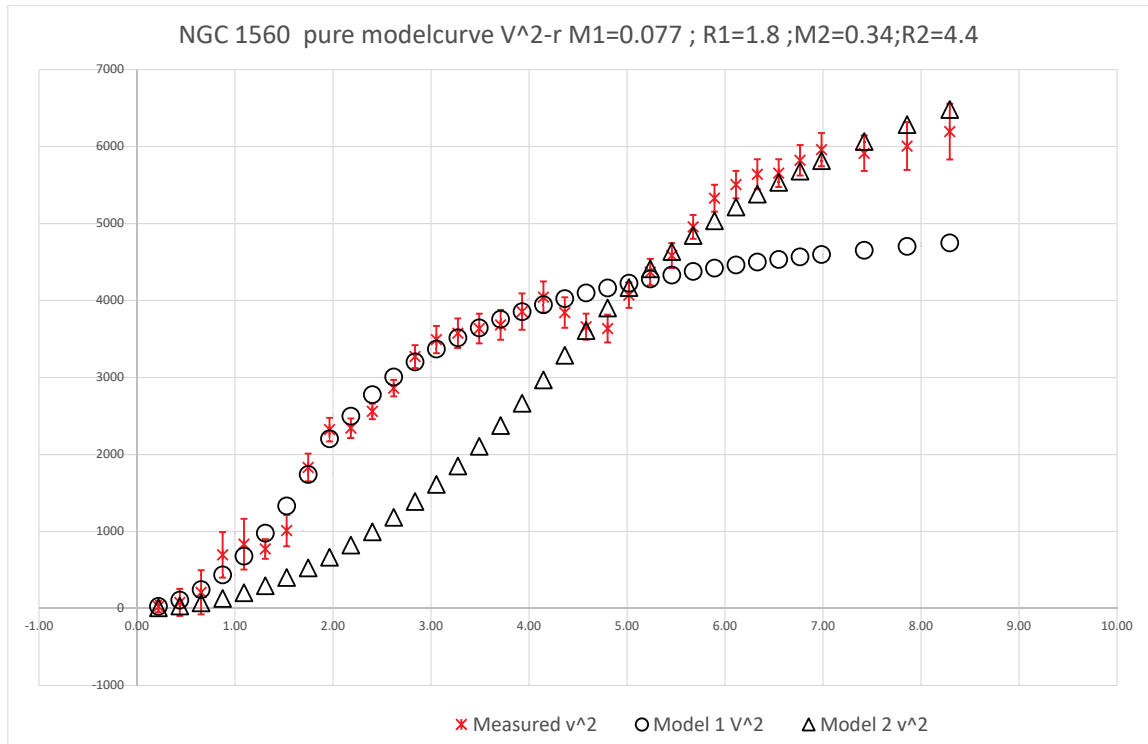


FIG. 4. Dual fit of NGC1560,  $V_{orb}^2$  against  $r$ .

While analyzing the SPARC database, this shift proved quite general. In the following I will give one example of such a shift, with galaxy DDO161. First I give the dual plot of the rotation curve, then I compare the result against the mass distribution made available by SPARC (from the [MassModels-LTG.zip](#) file).

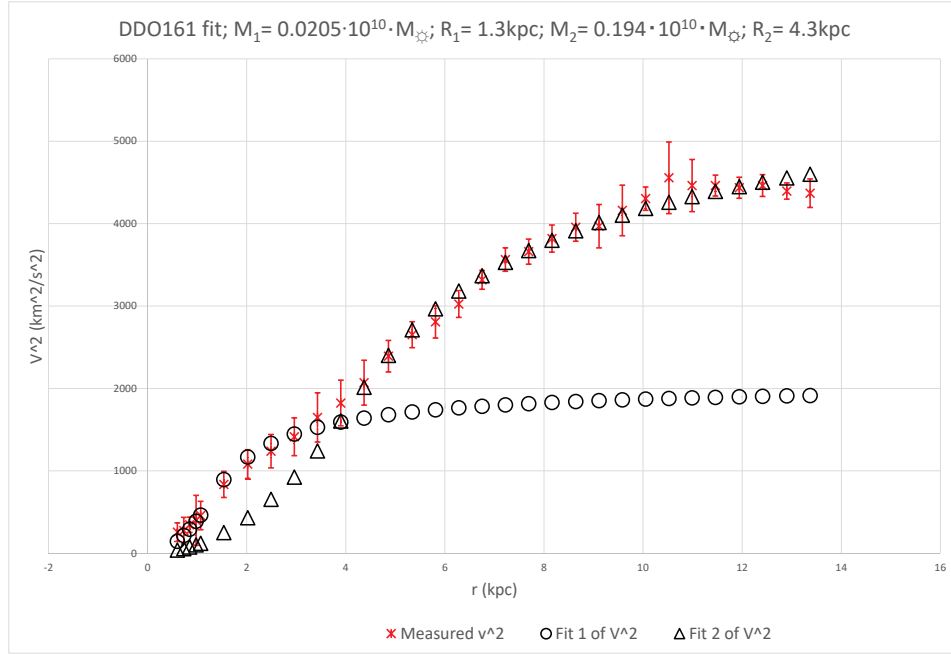


FIG. 5. DDO161,  $V_{orb}^2$  against  $r$ .

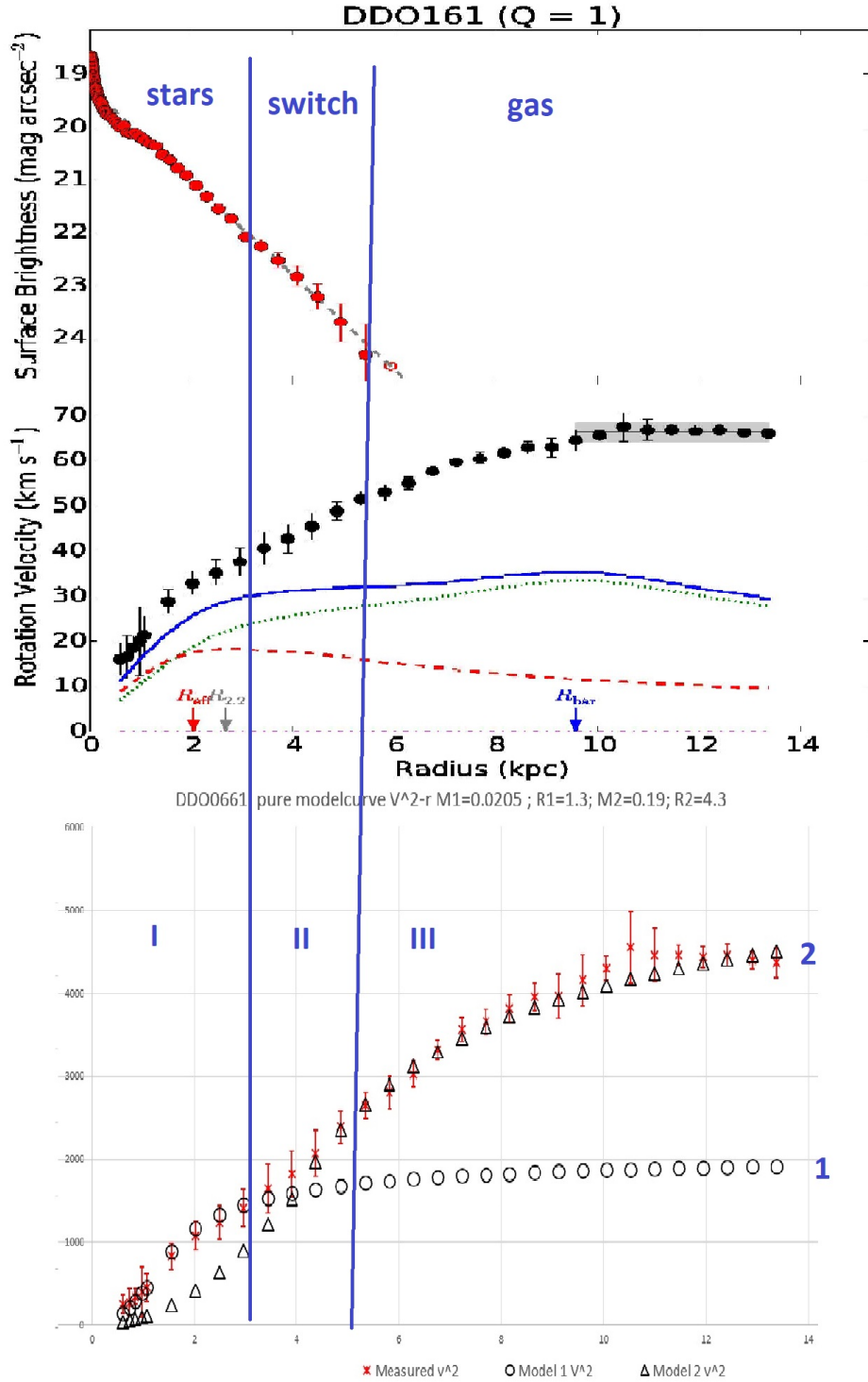


FIG. 6. DDO161,  $V_{orb}^2$  against  $r$ , and the stars and gas mass distribution plot. The red dots are the luminous stars, the dotted line the gas contribution to the mass and the red bars line represent the stars contribution to the velocity curved, assuming the virial theorem.

#### IV. THE SIXTEEN F-SERIES ROTATION CURVE FITS

In this section, I present the ‘constant Lagrangian’ rotation curve fits of the sixteen galaxies whose name starts with an F in of the [SPARC database](#), including the error margins, as provided by (Lelli et al., 2016). Most of these are dual fits. It is to the reader to compare the results with the mass distribution graphs of SPARC (from the [MassModels-LTG.zip](#) file).

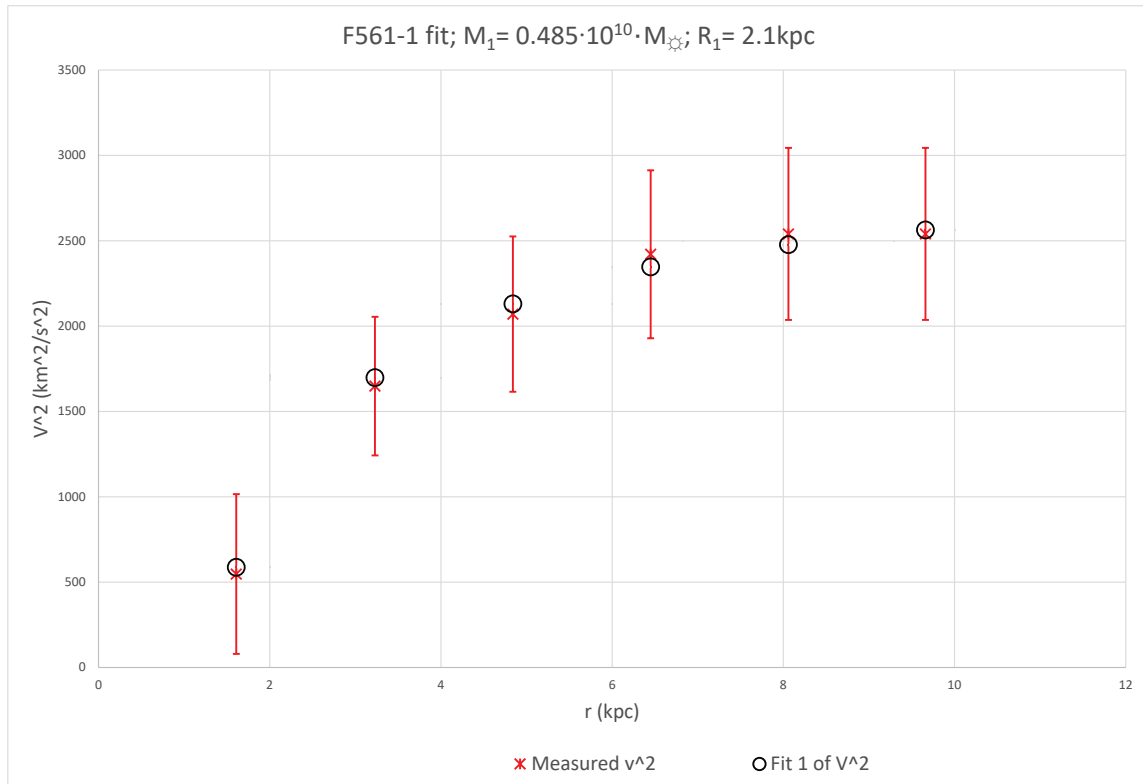


FIG. 7. F561-1,  $V_{orb}^2$  against  $r$

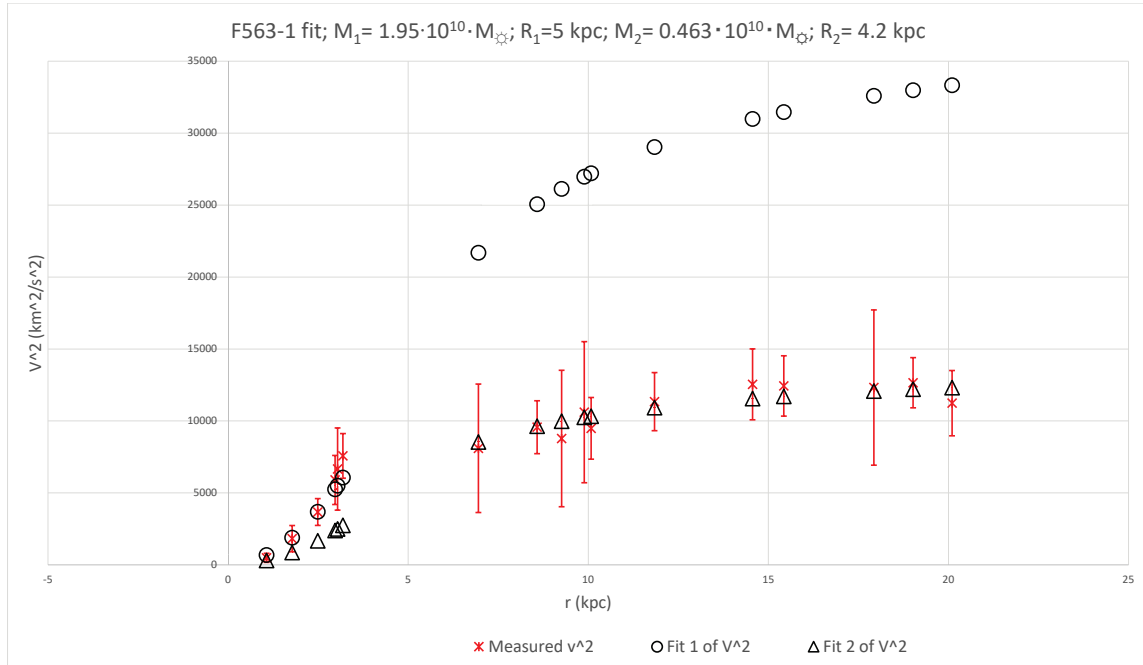


FIG. 8. F563-1,  $V_{orb}^2$  against  $r$

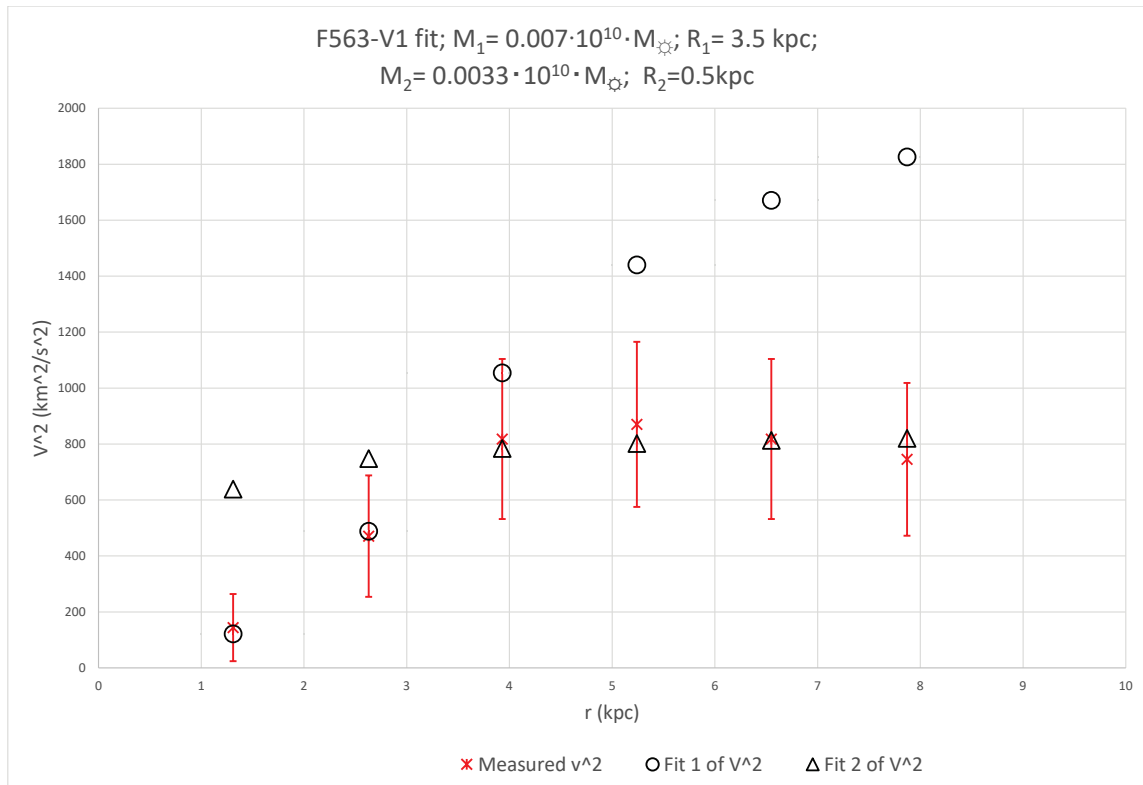


FIG. 9. F563-V1,  $V_{orb}^2$  against  $r$



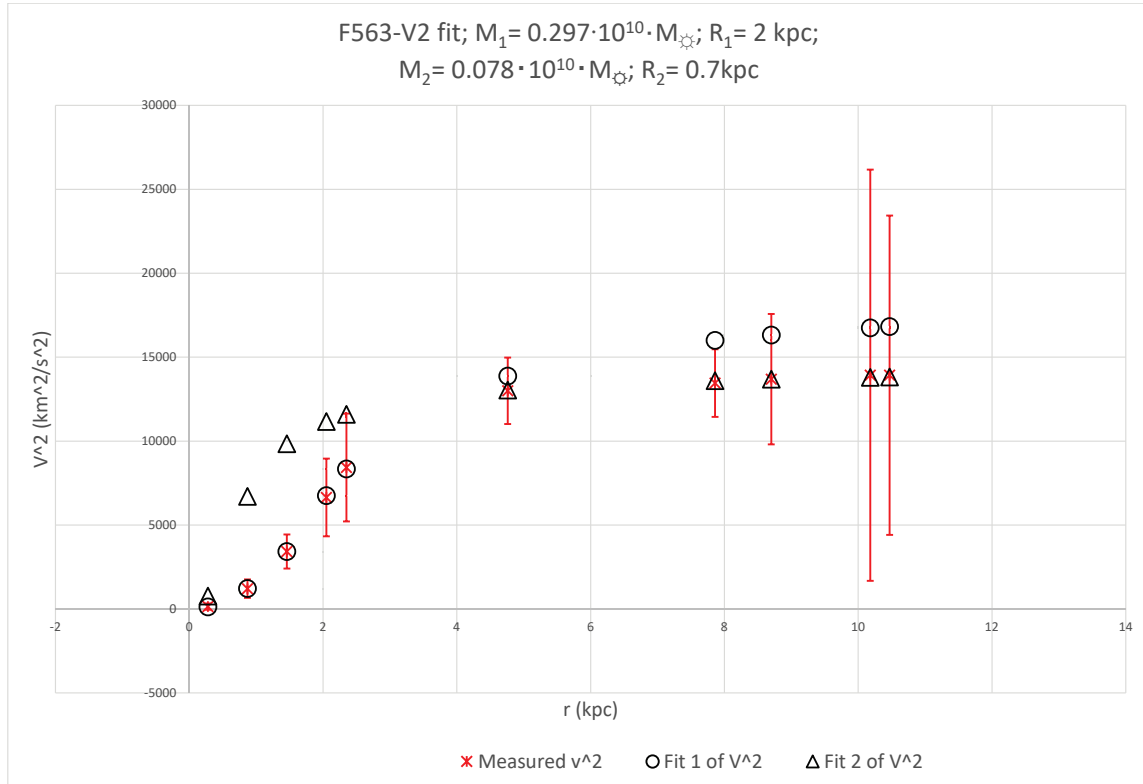


FIG. 10. F563-V2,  $V_{orb}^2$  against  $r$

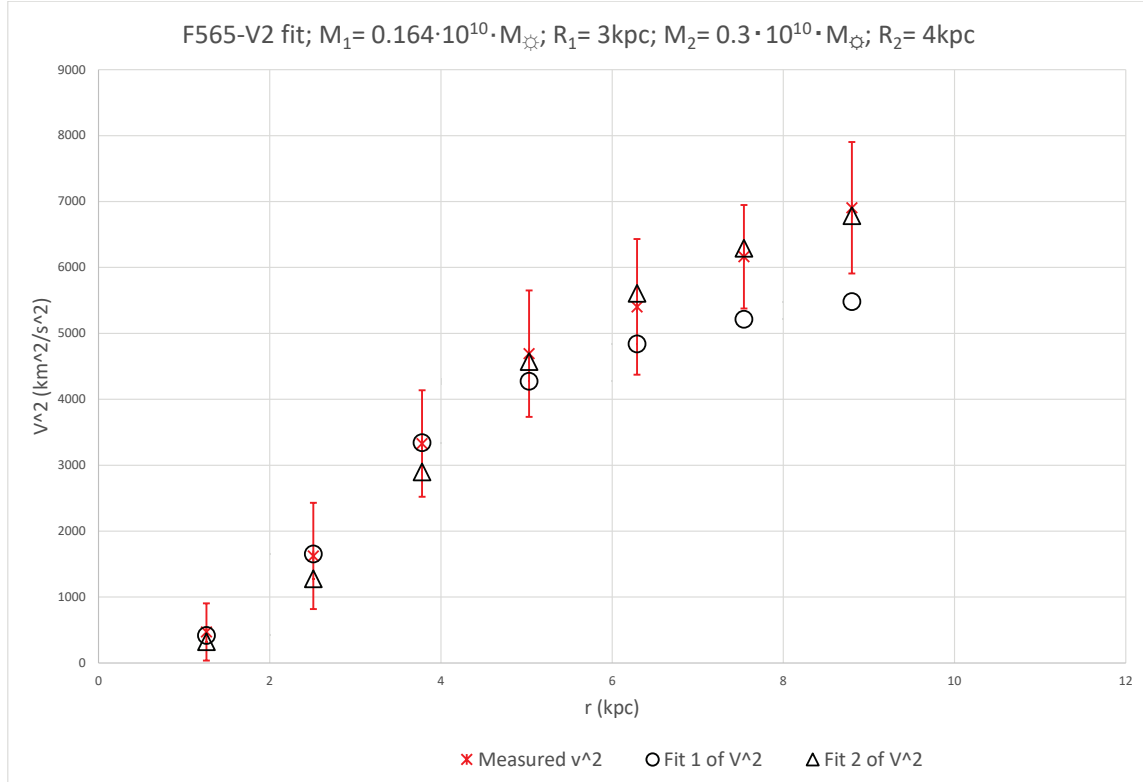


FIG. 11. F565-V2,  $V_{orb}^2$  against  $r$

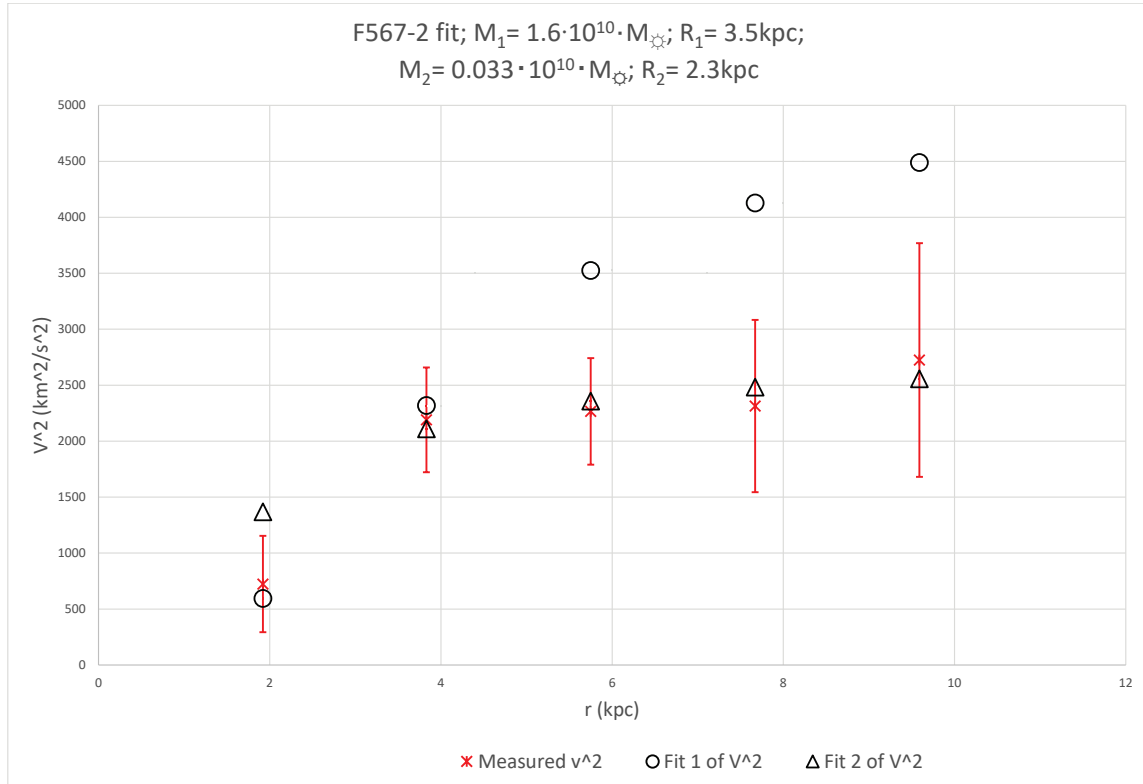


FIG. 12. F567-2,  $V_{orb}^2$  against  $r$

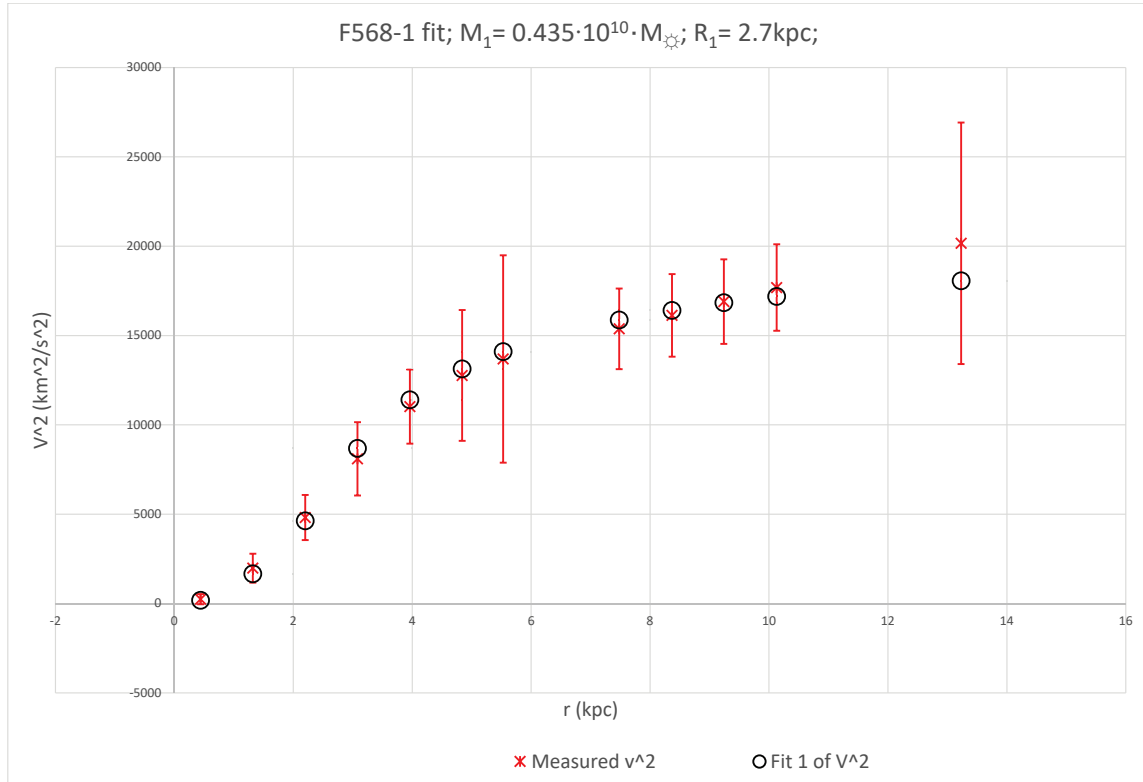


FIG. 13. F568-1,  $V_{orb}^2$  against  $r$

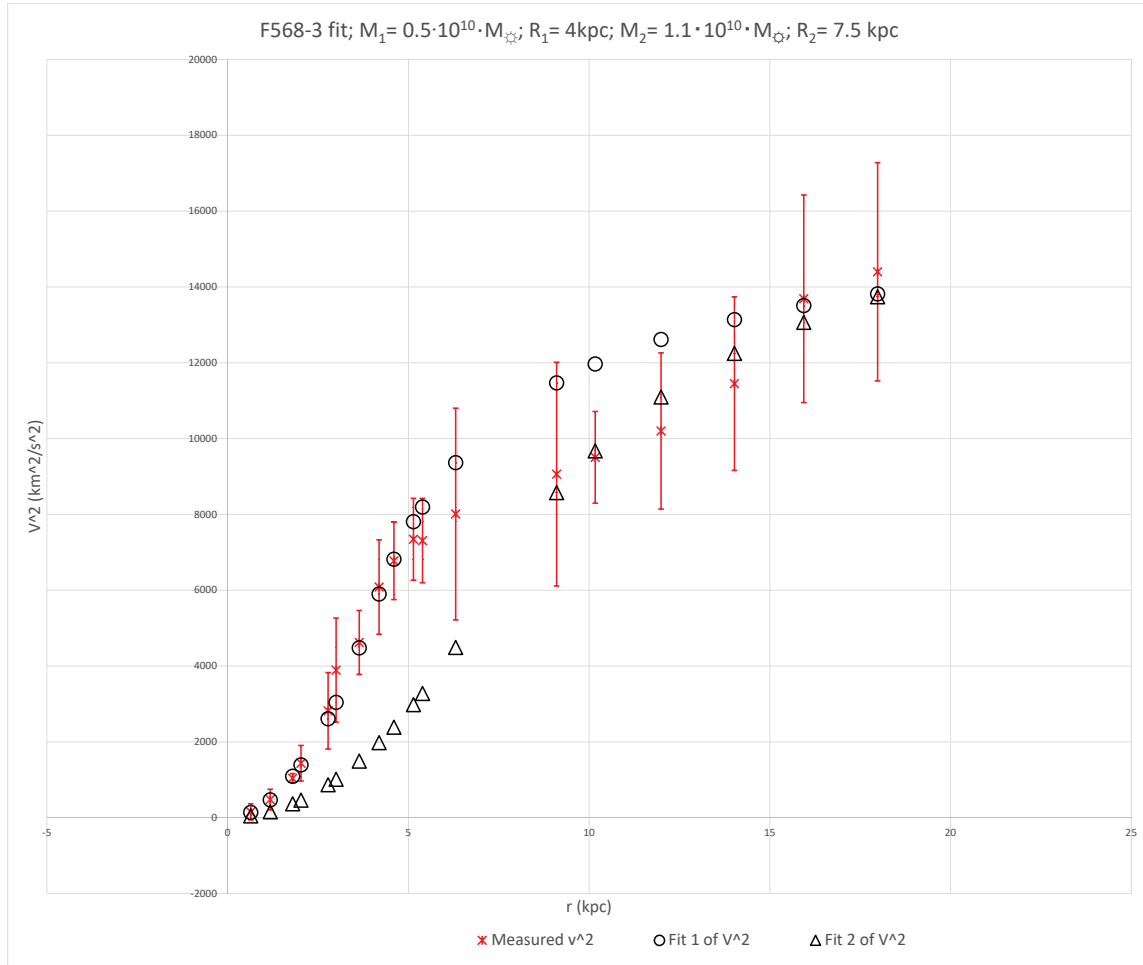


FIG. 14. F568-3,  $V_{orb}^2$  against  $r$

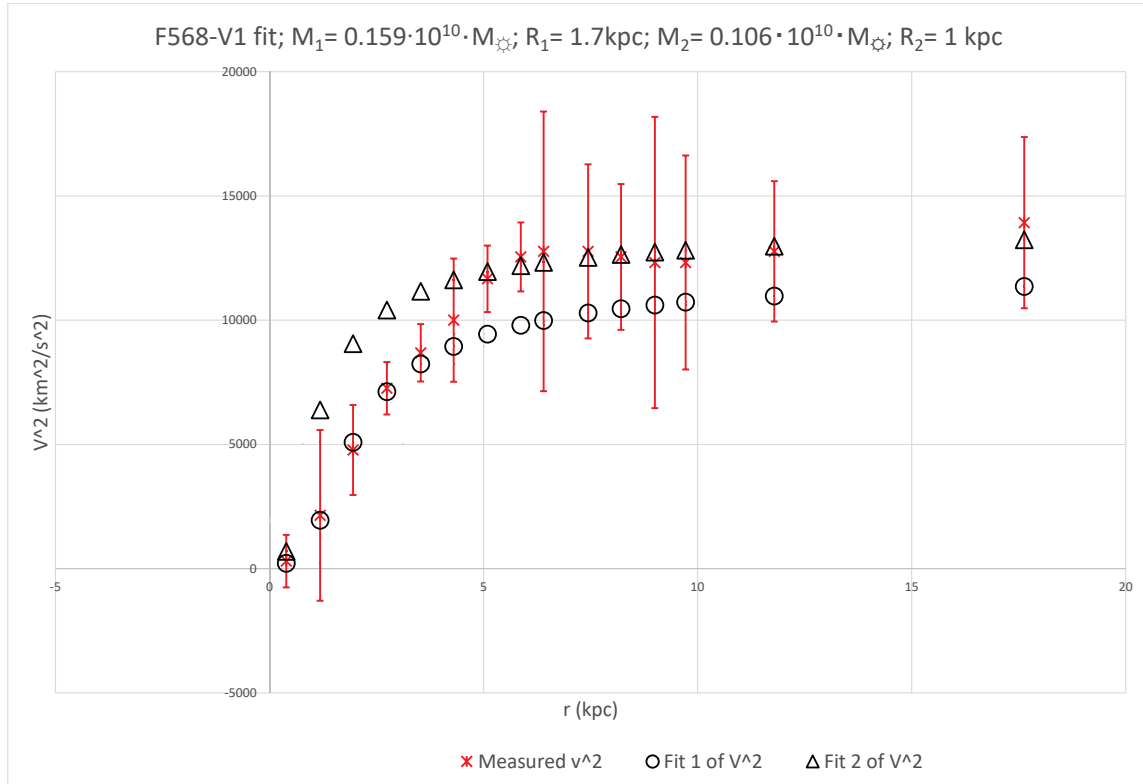


FIG. 15. F568-V1,  $V_{orb}^2$  against  $r$

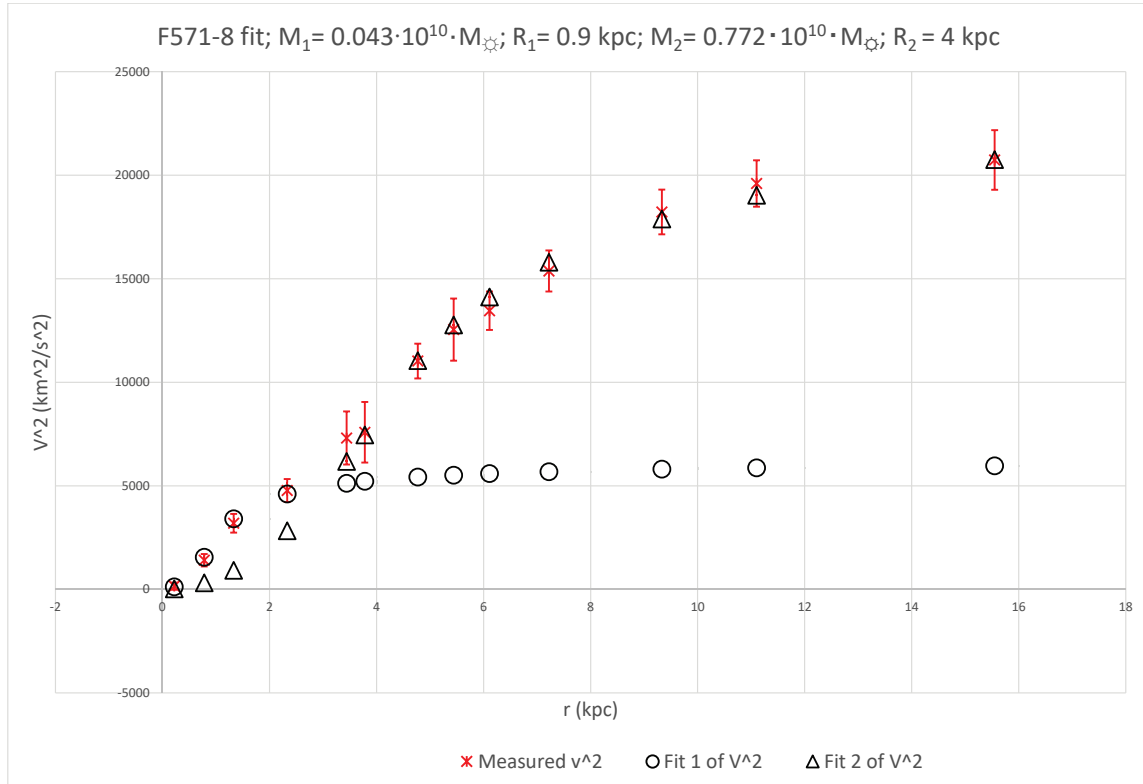


FIG. 16. F571-8,  $V_{orb}^2$  against  $r$

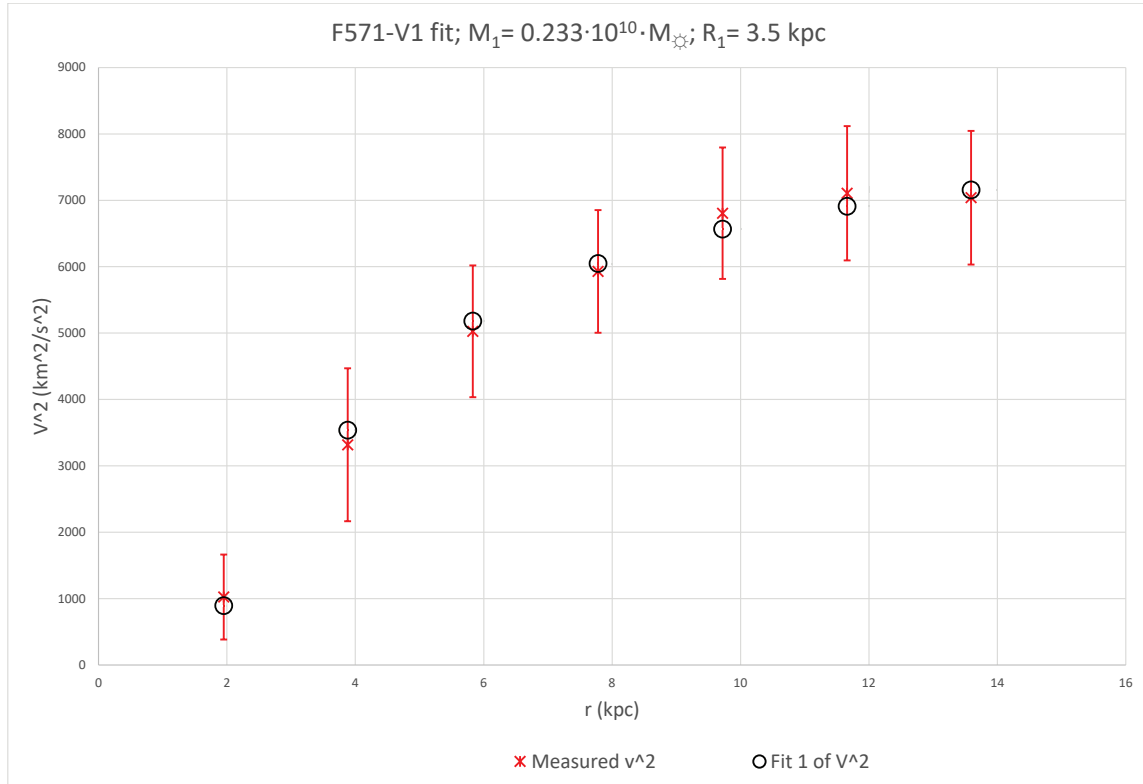


FIG. 17. F571-V1,  $V_{orb}^2$  against  $r$



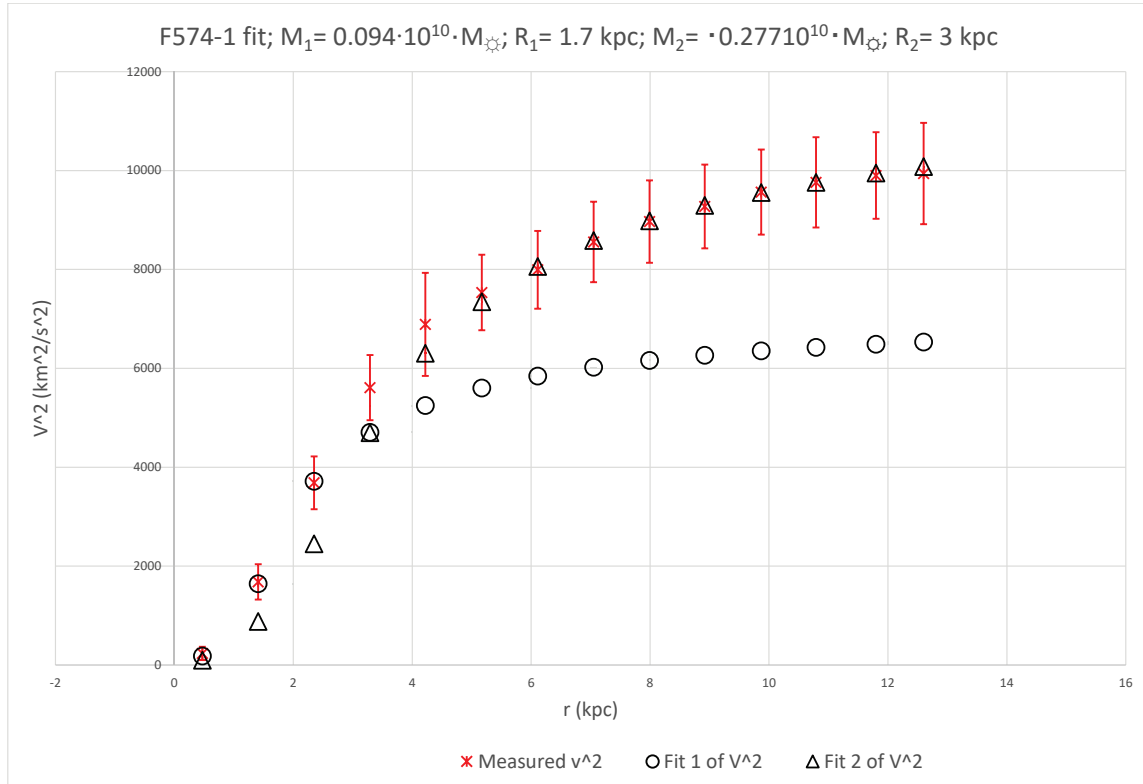


FIG. 18. F574-1,  $V_{orb}^2$  against  $r$

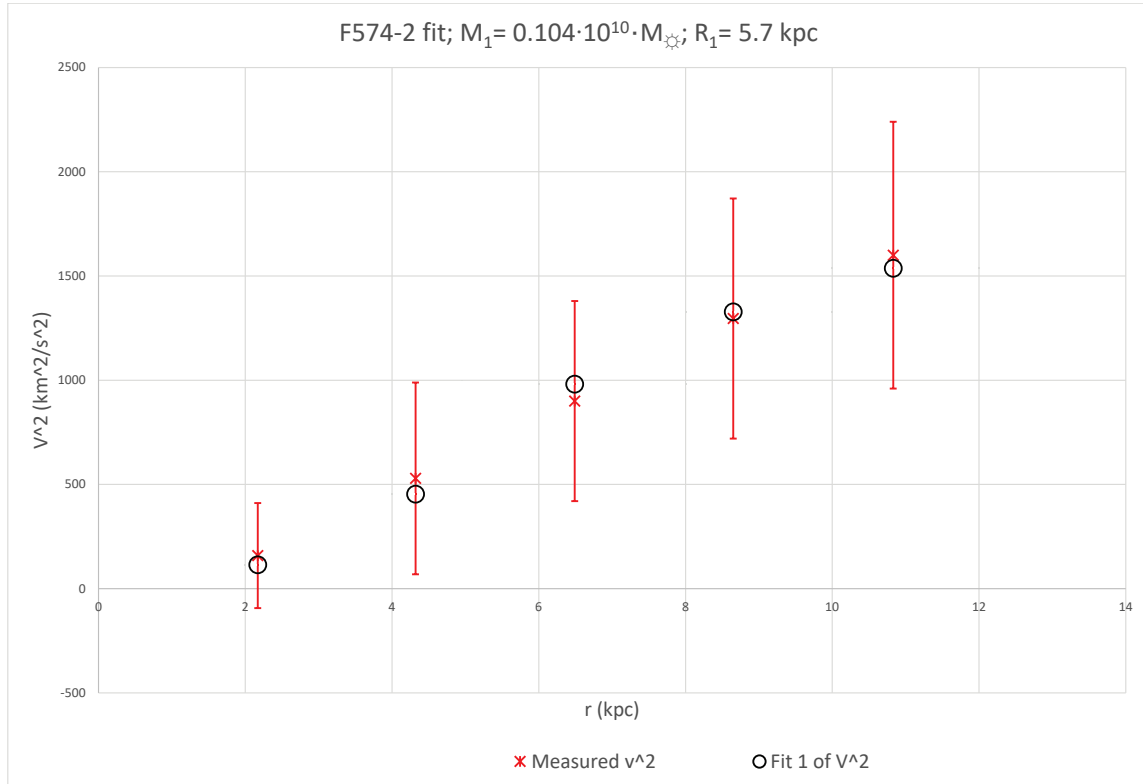


FIG. 19. F574-2,  $V_{orb}^2$  against  $r$

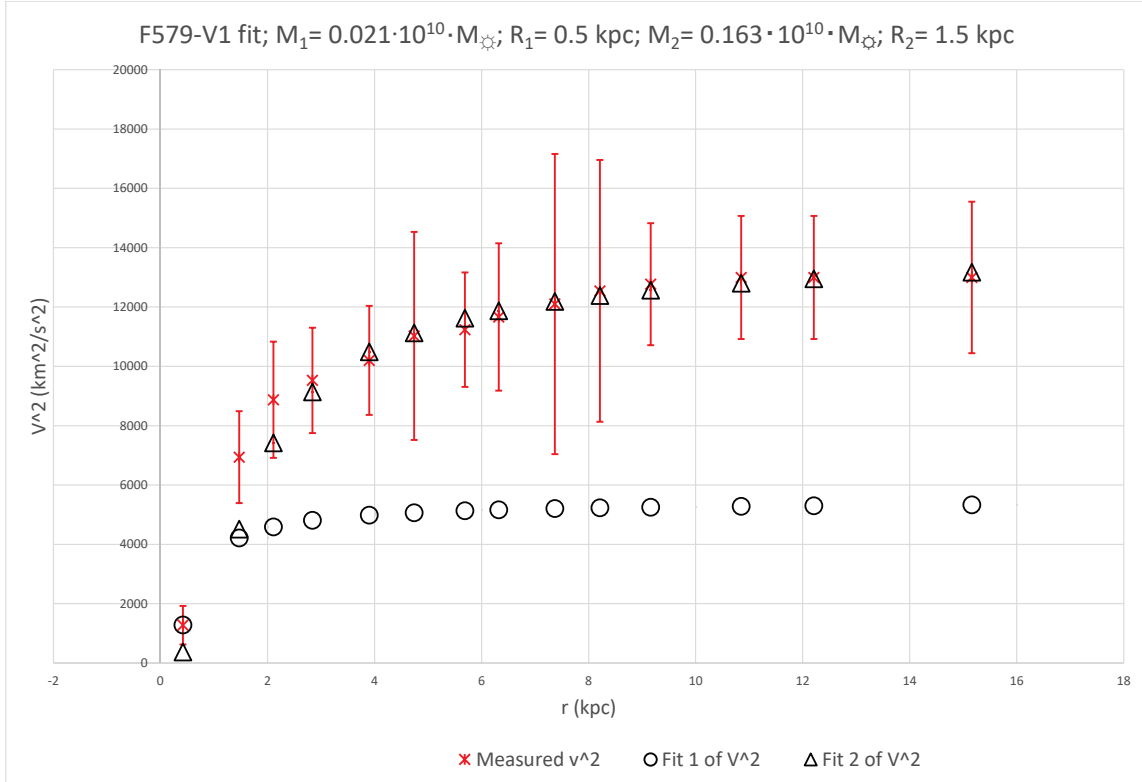


FIG. 20. F579-V1,  $V_{orb}^2$  against  $r$

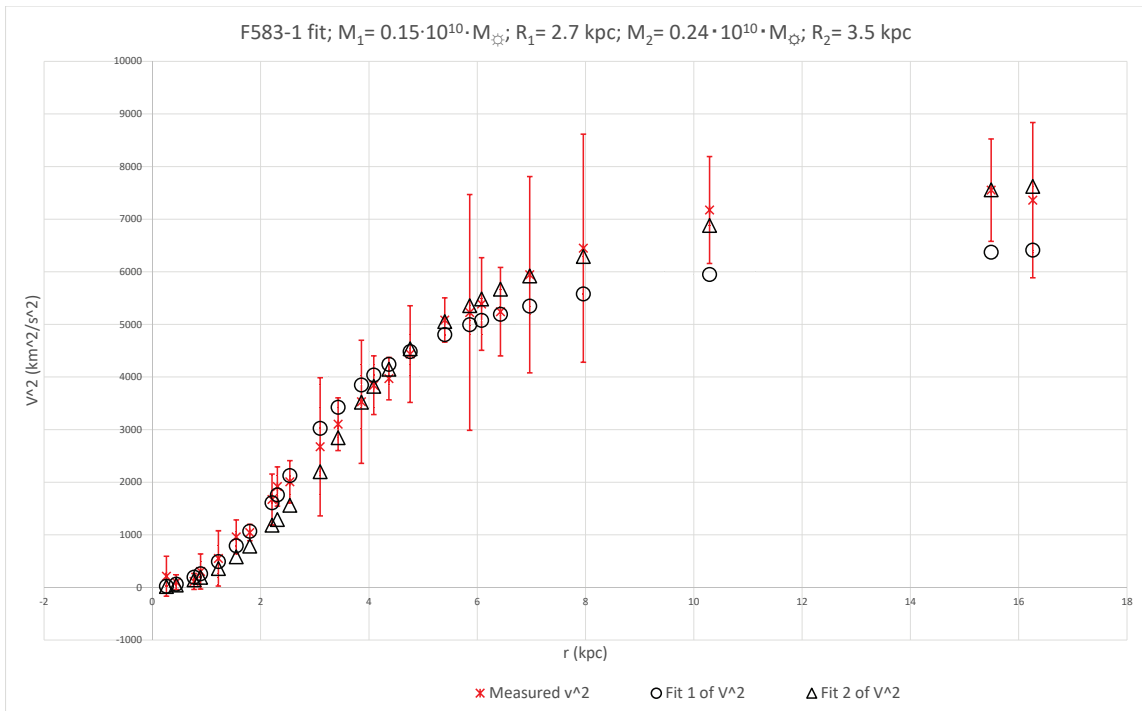


FIG. 21. F583-1,  $V_{orb}^2$  against  $r$

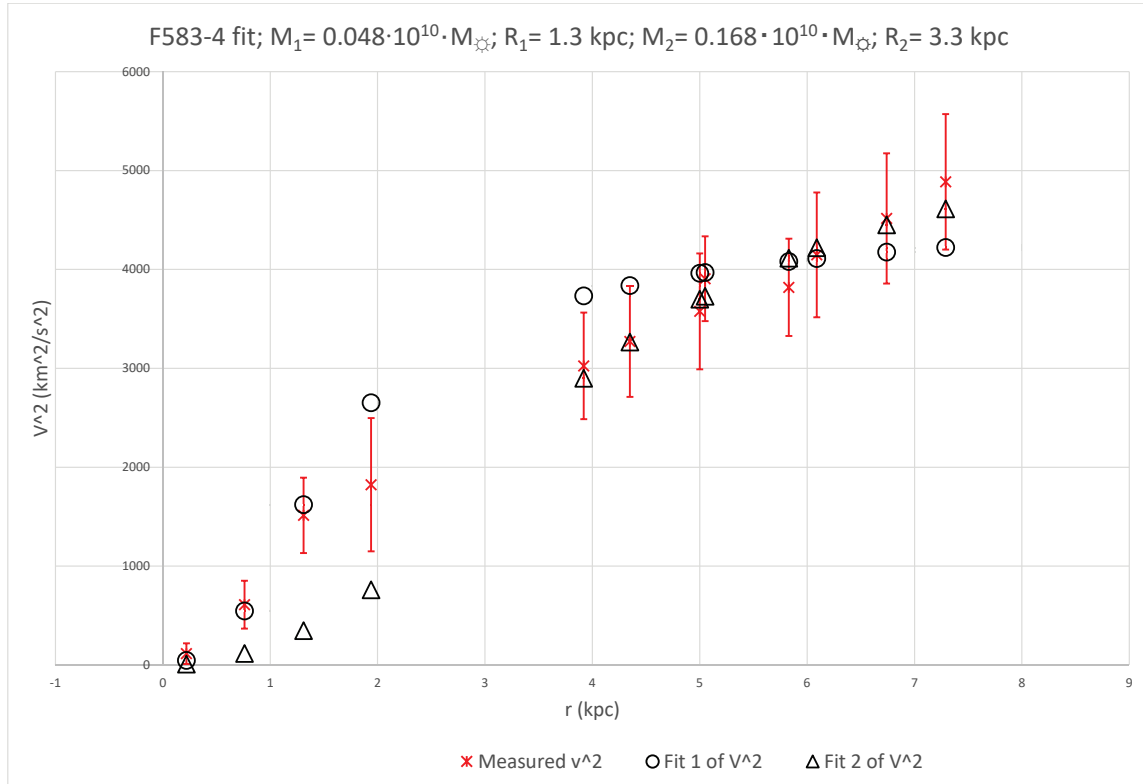


FIG. 22. F583-4,  $V_{orb}^2$  against  $r$

## V. CONCLUSION

Some of the sixteen rotation curves can be fit within the error margins with one model only. Most need a dual fit to remain within the error margins, with the shift from the one to the other as one goes through a distinct mass distribution shell. Some galaxies are more complicated.

In many cases, the shift from the first model to the second correspond with the shift from star dominated to gas dominated regions as given by the SPARC database. This correspondence happens so often that it cannot be a mere coincidence.

The fact that a correspondence exists between the shift of the fits of the rotation curves from the first model to the second and the mass distribution as composed of a bulge with a subsequent mass shell indicates a mass determined curvature of the metric. The ‘constant Lagrangian’ model is a time bubble halo model and as such represents a curvature of space-time in and around the model galaxy. In General Relativity, mass is the source of curvature. The ‘constant Lagrangian’ model should of course follow this basic axiom.

A follow-up questions should be if a three model fit would be needed for the rotation curve of a galaxy with a clear distinct threefold composition of bulge, disk and gas clouds. The question is if there is a predictive power in these models as to the mass and the radius of real galaxies.

Another follow-up research question is to check if or to what extend the  $(M, R)$  values of the model correspond to the measured mass and respective radius of those galaxies.

For the moment, the result of this paper with the fits of the F-series of the SPARC database is that the ‘constant Lagrangian’ model has promising powers.

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