#### How a Minimum time step and Formation of Initial Causal structure in space-time may void the Penrose Singularity theorem, as in Hawkings and Ellis's 1973 write ups

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#### Abstract

Using a root finder procedure to obtain  $\Delta t$  were we use an inflaton value due to use of a scale factor  $a \sim a_{\min} t^{\gamma}$  if we furthermore use  $\delta g_{ii} \sim a_{\min}^2 \cdot \phi_{initial}$  From use of the inflaton, we initiate a procedure for a minimum scale factor, which would entail the  $a_{\min}^4 < 0 \Rightarrow a_{\min} \sim \frac{\pm \sqrt{\#} \cdot (1+i)}{\sqrt{2}}$ , for a sufficiently well

placed frequency  $\omega$ . If the Non Linear Electrodynamics procedure of Camara et.al. of General relativity were used, plus the modified Heisenberg Uncertainty principle, of Beckwith, and others, is used, I,e.  $\Delta E \sim \left[ \hbar / \Delta t \cdot \left( \delta g_{tt} \sim a_{\min}^2 \cdot \phi_{initial} \right) \right] \Big|_{\text{Pre-Planckian}} \sim H(Hubble) \text{ we come due to a sufficiently high}$ frequency a case for which  $a_{\min} \sim \frac{\pm \sqrt{\#} \cdot (1+i)}{\sqrt{2}}$  implies a violation of the Penrose singularity theorem. I.e. this is in lieu of  $\Lambda_{initial} = \Lambda_{Today}$ . If this is not true, i.e. that the initial  $\Lambda \gg \Lambda_{Today}$ , then we will likely avoid  $a_{\min} \sim \frac{\pm \sqrt{\#} \cdot (1+i)}{\sqrt{2}}$  for reasons brought up in this manuscript.

Key words Inflaton physics, causal structure, non Linear Electrodynamics.

#### **I.** Framing the initial inquiry

Here, the idea would be, to make the following equivalence, namely look at, [1] as well as our own derivation

$$\left[\left[\frac{\Lambda_{Max}r^4}{8\pi G}\right] \cdot \left(4/3\right) \cdot \left[\frac{2\pi^2 g_*}{45}\right]^{1/3}\right]^{3/4} \sim S_{initial}$$
(1)

We furthermore, make the assumption of a minimum radius of [2,3]

$$R_{initial} \sim \frac{1}{\#} \ell_{Ng} < l_{Planck} \tag{2}$$

We will initially be assuming that the Cosmological constant remains as it is today, and not refer to the situation given in [4] as given by Park et. al, where the initial value of the cosmological constant could be much higher initially.

This Eq. (1) will be put as the minimum value of r, where we have in this situation [5, 6]

$$\#bits \sim \left[\frac{E}{\hbar} \cdot \frac{l}{c}\right]^{3/4} \approx \left[\frac{Mc^2}{\hbar} \cdot \frac{l}{c}\right]^{3/4}$$
(3)

And if M is the total space-time energy mass, for initial condition [5,6, where

$$S_{initial} \sim n_{graviton} \sim initial graviton count$$
 (4)

M then will be defined by the mass of a massive graviton [7], times, the graviton count, as given in (4) and the modified uncertainty principle, [3] and the Camarra et.al. defined Hubble parameter, given in [2]

$$\Delta E \sim \left[ \hbar / \Delta t \cdot \left( \delta g_{tt} \sim a_{\min}^2 \cdot \phi_{initial} \right) \right]_{\text{Pre-Planckian}}$$

$$\sim H(Hubble) \sim \sqrt{\frac{4\pi GB^2}{3c^2 \mu_0} \cdot \left( 1 - 8\mu_0 \omega B^2 \right) + \frac{\Lambda c^2}{3}}$$

$$\sim \sqrt{\frac{4\pi GB^2}{3c^2 \mu_0} \cdot \left( 1 - 8\mu_0 \omega B^2 \right) + \frac{\Lambda_{Today} c^2}{3}}$$
(5)

This will lead to

$$a_{\min} \sim \frac{\pm \sqrt{\#} \cdot (1+i)}{\sqrt{2}} \tag{6}$$

Whereas if  $\Lambda \gg \Lambda_{Today}$ , Eq. (6) likely will not hold, and we also state that Eq. (6) is a violation of the Penrose singularity theorem as written up in [8], whereas we also have that we are using the Padmanbhan results as given in [9,10, 11] to the effect that we are employing

$$a \sim a_{initial} t^{\gamma}$$

$$\phi = \sqrt{\frac{\gamma}{4\pi G}} \ln \left[ \sqrt{\frac{8\pi V_0 G}{\gamma (3\gamma - 1)}} t \right]$$
(7)

While adhering to a potential in line with

$$V = V_0 \exp\left[\left\{-\sqrt{\frac{16\pi G}{\gamma}}\right\} \cdot \phi(t)\right]$$
(8)

We next then go to the results given in [9] which have been become an approved for publication by JHEPGC.

## **1.** Examination of the minimum time step, in Pre-Planckian Space-time as a Root of a Polynomial Equation.

We initiate our work, citing [9] to the effect that we have a polynomial equation for the formation of a root finding procedure for  $\Delta t$ , namely if

$$\Delta t \cdot \left[ \left( \sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1 \right) - \frac{\left( \sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1 \right)^2}{2} + \frac{\left( \sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1 \right)^3}{3} - \dots \right]$$

$$\approx \left( \sqrt{\frac{\gamma}{\pi G}} \right)^{-1} \frac{48\pi \hbar}{a_{\min}^2} \cdot \Lambda$$
(9)

From here, we then cited, in [9], using [6] a criteria as to formation of entropy, i.e. If  $\Lambda$  is an invariant cosmological 'constant' and if Eq. (10) holds, we can use the existence of nonzero initial entropy as the formation point of an arrow of time.

$$S_{\Lambda}|_{Arrow-of-time} = \pi \cdot \left(\frac{R_c|_{initial} \sim c \cdot \Delta t}{l_{Planck}}\right)^2 \neq 0$$
(10)

This leads to the following, namely in [9] we make our treatment of the existence of causal structure, as given by writing its emergence as contingent upon having

$$\left(\frac{R_c|_{initial} \sim c \cdot \Delta t}{l_{planck}}\right) \sim \mathcal{G}(1)$$
(11)

The rest of this article will be contingent upon making the following assumptions. FTR

That we will drop most of the terms in the expansion of Eq. (9) and instead of a huge infinite expansion of terms, pick instead using

$$\Delta t \cdot \left(\sqrt{\frac{8\pi G V_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1\right) \approx \left(\sqrt{\frac{\gamma}{\pi G}}\right)^{-1} \frac{48\pi\hbar}{a_{\min}^2 \cdot \Lambda}$$
(12)

This is assuming here that the terms in  $\left(\sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1\right)$  are extremely small, which permits us to

come up with a quadratic expression of the term  $\Delta t$  which is of course useful as to what we do next, i.e.

If we make use of the Peebles relationship [12] of what would be occurring just before and at the start of expansion of the universe, i.e. the causal structure as given by [9] as, using the Keiffer result of [13] so as to get

$$g_{tt} \sim \delta g_{tt} \approx a_{\min}^{2} \phi_{initial} \ll 1$$

$$\xrightarrow{Pre-Planck \rightarrow Planck} \delta g_{tt} \approx a_{\min}^{2} \phi_{Planck} \sim 1$$

$$\Leftrightarrow \left( \frac{R_{c}|_{initial} \sim c \cdot \Delta t}{l_{Planck}} \right) \sim \mathcal{G}(1) \Big|_{Planck}$$
(13)

#### 2. CONSEQUENCES, In terms of the minimum scale factor

We then use the Peebles result [12] for the strain of space-time at the START of expansion result of

$$\Delta E \sim \left[ \hbar / \Delta t \cdot \left( \delta g_{tt} \sim a_{\min}^{2} \cdot \phi_{initial} \right) \right]_{\Pr e-Planckian}$$

$$\sim H(Hubble) \sim \sqrt{\frac{4\pi GB^{2}}{3c^{2}\mu_{0}} \cdot \left(1 - 8\mu_{0}\omega B^{2}\right) + \frac{\Lambda c^{2}}{3}}$$
(14)
$$\sim \sqrt{\frac{4\pi GB^{2}}{3c^{2}\mu_{0}} \cdot \left(1 - 8\mu_{0}\omega B^{2}\right) + \frac{\Lambda_{Today}c^{2}}{3}}$$

The key result is that we have a quadratic expression for the  $\Delta t$  term, as indicated by (12) with the result that there is a solvable expression in terms of  $\Delta t$ , so that then, we can take the square of the terms of Eq. (14) with using the expression of Eq. (7) above, in order to obtain after using an expansion of Ln x, (if 0<x<2) from [14] to get, then, after algebra

$$\frac{\gamma}{4\pi G} \left( \sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1 \right)^2$$

$$\sim \frac{\hbar^2}{a_{\min}^4 \cdot (\Delta t)^2 \cdot \left( \frac{4\pi GB^2}{3c^2 \mu_0} \cdot (1 - 8\mu_0 \omega B^2) + \frac{\Lambda_{Today} c^2}{3} \right)}$$
(15)

#### 3. Conclusion, two parts

#### **3a.** So what if the Denominator of Eq. (15) is less than Zero?

If that happens, due to either a strange, very high frequency value, and a small cosmological constant, we then have

$$a_{\min}^4 < 0 \Longrightarrow a_{\min} \sim \frac{\pm \sqrt{\# \cdot (1+i)}}{\sqrt{2}}$$
 (16)

Note here that when this happens, we have two equally admissible solutions for the scale factor, minimum, and the consequences if # is a real number, that then we have a contradiction with what is called Theorem 3, Hawking (1967) as cited on page 271, of [8] we have that

Theorem 3: If  $R_{ab}K^aK^b \ge 0$  for every non space-like Vector K

- a. The strong casuality condition holds on  $\left( ilde{M}, g 
  ight)$ ,
- b. There is some past –directed unit timelike vector W at a point p, and a positive constant b, such that if V is the Unit tangent fector to the past directed timelike geodestic through p, then on each geodestic the expansion  $\theta \equiv V_{;a}^{a}$  of these gebnodestics becomes less than -3c/p, within a distance b/c from p. Where  $c = -W^{a}V_{a}$ , i.e. then there is a past incomplete non space-like geodesic through p.

One does not have a curve violating the casuality conditions as is asserted by Hawkings and Ellis, 1973. I.e. there is, if this occurs at the causal boundary, instead, a bifurcation point at the surface of the causal set, with real and imaginary components, but the incompleteness of the non space geodestic through a point p, if it is on the surface of the causal surface, as defined by Eq. (13) is not due to a point p-. It is well known that certain Kerr black hole models, as in page 465 of Ohanian and Ruffini [14] involve the use of  $g_{tt} \sim 0$  for their horizon surfaces and the definition of a 'plate disc singularity surface but we are instead employing,  $\delta g_{tt} \sim a_{\min}^2 \cdot \phi_{initial}$ 

I.e. precisely because we have avoided using  $g_{tt} \sim 0$  as was done in the Kerr black holes, as given in [14] but instead have the  $\delta g_{tt} \sim a_{\min}^2 \cdot \phi_{initial}$  plus the situation we wish to avoid, that of instead looking at  $a_{\min}^4 < 0 \Rightarrow a_{\min} \sim \frac{\pm \sqrt{\#} \cdot (1+i)}{\sqrt{2}}$ , that a causal surface, would be formed on a sphere of space time

which would in itself violate the 3<sup>rd</sup> Penrose theorem

### **3b.** So what happens if $\Lambda_{initial} >> \Lambda_{Today}$ ?

The second case to consider would be if we have, instead of today's version of the cosmological constant, a large valued initial cosmological constant, in which then

$$\frac{\gamma}{4\pi G} \left( \sqrt{\frac{8\pi GV_0}{\gamma \cdot (3\gamma - 1)}} \cdot \Delta t - 1 \right)^2 \sim \frac{\hbar^2}{a_{\min}^4 \cdot (\Delta t)^2 \cdot \left(\frac{\Lambda_{initial}c^2}{3}\right)} \\ \& \qquad (17) \\ \Lambda_{initial} >> \Lambda_{Today}$$

We argue that then, there is no reason for assigning a singularity, but it would in line with [4]

, i.e. assigning an almost infinite value for the initial cosmological constant.

Different variants of the above can be imagined, and of course one should be considering [16] in the reformulation of the Causal structure boundary idea. In addition the points brought up as to [17] to [21] of the nonlinear electrodynamics cosmology should be utilized as a refinement as to the Hubble parameter as outlined in Eq. (5) above

# **3c.** Otonion geometry and non-commutativity as a future project to be combined with our present inquiry?

We should close with one reference as to the Octonionic geometry program as follows, We may be seeing instead of just our roof finder iterations, as outlined above, an exploration into non commutative geometry. This is what I am referring to, and it is from [22].

From [22]

Quote:

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i.e.

. The change in geometry is occurring when we have first a pre quantum space time state, in which, in commutation relations [23] (Crowell, 2005) in the pre Octonion space time regime no approach to QM commutations is possible as seen by.

$$[x_{j}, p_{i}] \neq -\beta \cdot (l_{Planck} / l) \cdot \hbar T_{ijk} x_{k} \quad and \quad does \quad not \rightarrow i\hbar \delta_{i,j}$$
(18)

Eq. (18) is such that even if one is in flat Euclidian space, and i= j, then

$$\left[x_{j}, p_{j}\right] \neq i \cdot \hbar \tag{19}$$

In the situation when we approach quantum "octonion gravity applicable" geometry, Eq.(18) becomes

$$[x_{j}, p_{i}] = -\beta \cdot (l_{Planck} / l) \cdot \hbar T_{ijk} x_{k} \xrightarrow{Approaching-flat-space} i\hbar \delta_{i,j}$$
(20)

End of quote

We assert that the issues as of Eq. (18) to Eq. (20) if done in higher dimensional analogues, taking into account non commutative initial geometry and the approach to communitive geometry as outlined in [23] in time, if twinned directly with an analysis of Eq. (15) to Eq. (17) may in time help us delineate the future of space time research in the early universe.

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