

Black hole as a gas of strings

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ABSTRACT- We used the micro canonical ensemble in order to evaluate the entropy of a gas of strings confined by its Schwarzschild radius. The results are compared with the entropy of the gravitational field calculated by Rothman and Anninos and with the well-known Bekenstein-Hawking entropy of a black hole.

1 - INTRODUCTION

Rothman and Anninos [1] used the relation $S = k_b \ln \Omega$ as a means to evaluate the entropy S of the gravitational field, where Ω is the phase space of the field bounded by a Hamiltonian. In their work, the phase space is calculated for gravitational waves and radiation and density perturbations in expanding FLRW spacetimes, attributing entropy to a lack of knowledge in the exact field configuration.

Meanwhile we may think of a universe populated by a gas of strings, and perhaps would be interesting to evaluate the entropy of this gas in order to compare with Rothman and Anninos calculations.

2 - THE HAMILTONIAN OF THE GAS OF STRINGS AND PHASE SPACE CALCULATION

The Hamiltonian of the gas of strings can be written as

$$H = \sum_{i=1}^N \{c |\mathbf{p}_i| + a |\mathbf{r}_i|\}, \quad (1)$$

where c is the light speed in vacuum and a is the string tension and $(\mathbf{p}_i, \mathbf{r}_i)$ are labels locating the i -particle in the space-phase. The equation for the one-dimensional space-phase reads

$$(c |\mathbf{p}_i|)/E + (a |\mathbf{r}_i|)/E = 1. \quad (2)$$

The space of phase area (please see figure 1) is given by

$$A = (2 E^2)/(ca). \quad (3)$$

The hypercube volume in N -dimensions reads

$$V_N = A^N = \{2/(ca)\}^N E^{2N}. \quad (4)$$

Figure 1 (below), shows a graph of the trajectory of a particle in the space of phase. Vertical axis gives the momentum coordinate p measured in terms of E/c , and horizontal axis gives the position coordinate r in terms of E/a . We choose arbitrary values for the vertical and horizontal intercepts, namely $E/c = \pm 1$, and $E/a = \pm 2$.

Graph for $(-0.5)*\text{abs}(x)+1$, $0.5*\text{abs}(x)-1$

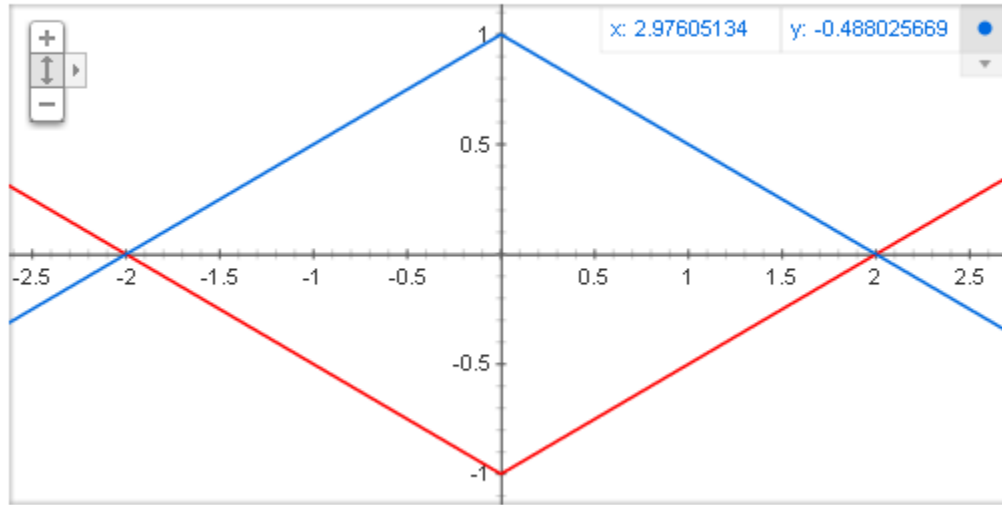


Figure 1- The 1-d space of phase

3 – BOHR-QUANTIZATION OF THE STRING HAMILTONIAN

Let us take

$$H = c p + a r, \quad (5)$$

and the rule for Bohr-quantization

$$r p = n \hbar, \quad n = 1, 2, 3, \dots, N. \quad (6)$$

Inserting (6) into (5) we obtain

$$H = c p + (n a \hbar) / p. \quad (7)$$

Next we take the minimum of (7), performing the calculation $\partial H / \partial p = 0$, and find

$$p_n = [(n a \hbar) / c]^{1/2}. \quad (8)$$

Putting (8) into (6) leads to

$$E_n = H(p_n) = 2 (a \hbar c)^{1/2} n^{1/2}. \quad (9)$$

Now if we denote N the maximum quantum number of the gas of strings, we get

$$E_N = 2 (a \hbar c)^{1/2} N^{1/2}, \quad \text{and} \quad (E_N)^{2N} = 4^N (a \hbar c)^N N^N. \quad (10)$$

Inserting (10) into (4), we obtain for the volume of phase space the relation

$$V_N = (16 \pi)^N \hbar^N N^N. \quad (11)$$

The next step is to use the recipe to evaluate Ω , working with the micro canonical ensemble (please see reference [2]). We have

$$\Omega = V_N / (\hbar^N N!) = \{[(16\pi)^N N^N] / N!\}. \quad (12)$$

4 - CALCULATION OF THE ENTROPY

As a means to obtain the entropy S of the gas of strings, we use the Boltzmann recipe, namely

$$S = k_B \ln \Omega = N k_B \ln(16 \pi e). \quad (13)$$

Figure 2 shows a graph of the string potential $U(r) = a |r|$, and the energy constant Mc^2 .

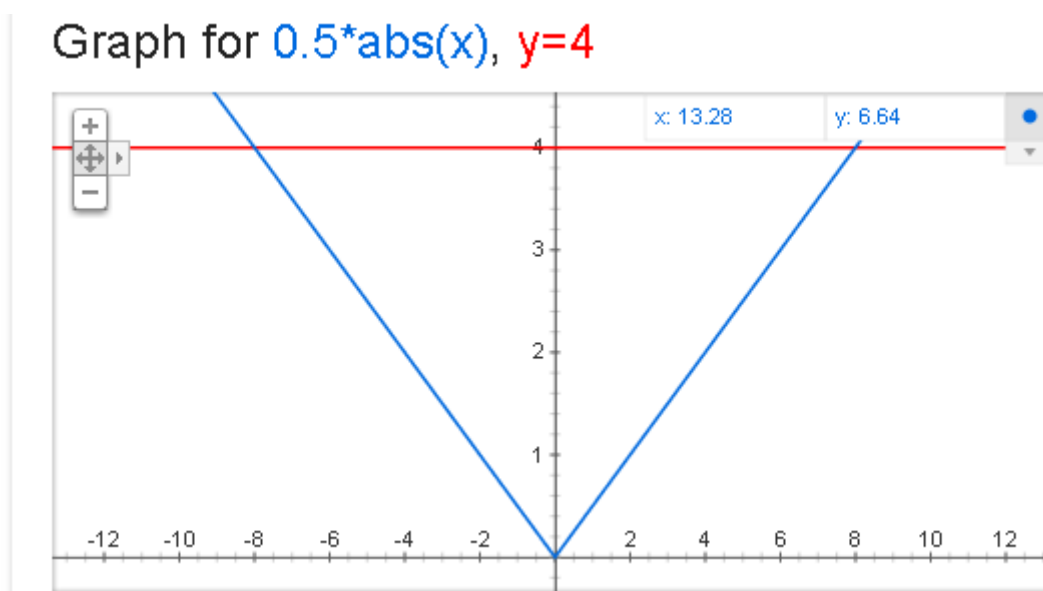


Figure 2 - The string potential $U(r) = a |r|$ (blue) and energy constant Mc^2 (red) are exhibited with arbitrary values of parameters, namely $a = .5$, $Mc^2 = 4$ and r_N (the classical turning point)=8.

The acknowledgment of the gas of strings as a black hole is done by making the identification the turning point r_N with the radius of Schwarzschild R_S of a black hole of mass M . Doing this, we have

$$r_N = N^{1/2} [(\hbar c)/a]^{1/2} = (2GM)/c^2 \equiv R_S. \quad (14)$$

Solving (14) for N , we obtain

$$N = (4a G^2 M^2) / (\hbar c^5). \quad (15)$$

Meanwhile the turning point condition reads

$$a r_n = M c^2, \text{ with } r_n \equiv R_S = (2GM)/c^2. \quad (16)$$

Relation (16) implies that

$$(2 a G)/c^4 = 1. \quad (17)$$

Inserting (17) into (15) we obtain

$$N = 2 (M/M_{\text{Pl}})^2, \quad (18)$$

where $M_{\text{Pl}} = (\hbar c/G)^{1/2}$ stands for the Planck's mass.

Finally by using (18) in (13) we get the entropy for this gas of strings confined in its Schwarzschild radius, namely

$$S = k_B [2 \ln(16\pi e)] (M/M_{\text{Pl}})^2. \quad (19)$$

5- CONCLUDING REMARKS

The entropy of the confined gas of strings given by (19) must be compared with

$S_{\text{R-A}} = k_B [4 \ln(2\pi e)] (M/M_{\text{Pl}})^2$, obtained by Rothman and Anninos[1] and with

$S_{\text{B-H}} = k_B [8 \pi^2] (M/M_{\text{Pl}})^2$, the well known result of Bekenstein-Hawking [3,4].

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