

# A Newtonian Model for Spiral Galaxy Rotation Curves

Geoffrey M. Williams  
email : gmwill@charter.net

## Abstract

A spiral galaxy is modeled as a thin, flat, axisymmetric disk comprising a series of concentric, coplanar rings. Using conventional Newtonian gravitation kinematics, it is shown that relatively flat velocity curves are produced by a variety of possible mass distributions in the disk. No halo of “dark matter” is needed to produce these rotation curves. Compared with a point mass at the center, the disk gravitational force grows with increasing distance from the disk center, crests and then slowly subsides beyond the disk perimeter. The model is applied to the NGC 3198, M31 and NGC 4736 galaxies, with ring masses adjusted to match the respective velocity profiles. Gravitational force fields in the disk are calculated, leading to direct estimates of enclosed galaxy mass. The mass distributions of several other spiral galaxies are analyzed, and their basic characteristics are charted in Appendix 2.

*Subject headings:* galaxies: kinematics and dynamics-gravitation-mass distribution

## Introduction

Spiral galaxies are disks containing billions of stars and matter, rotating about the disk center. This mass of material is held together by gravitational forces counteracting the centrifugal forces of disk rotation. When astronomers started reviewing the rotation curves of spiral galaxies, they were surprised to find, that as distance from the disk center increased, circular velocities did not decline.

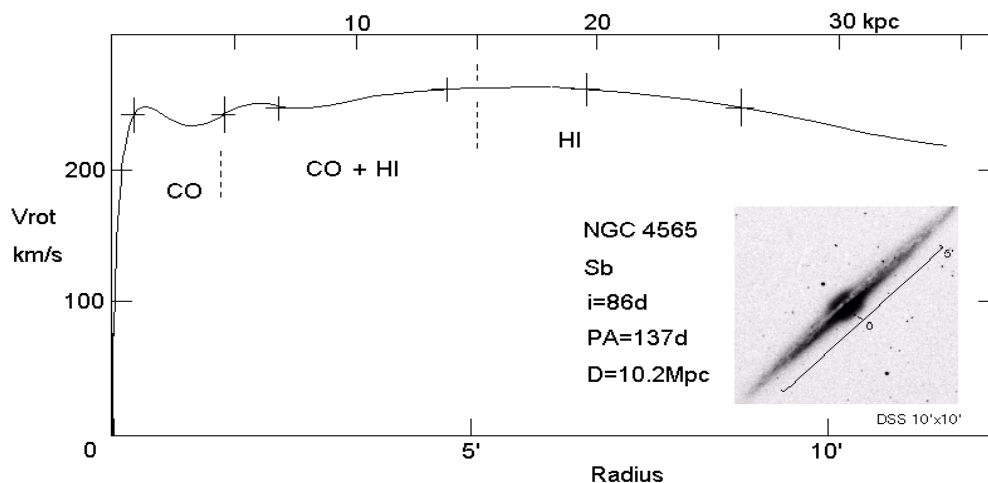


Figure 1. Rotation Curve for Spiral Galaxy NGC 4565 (Reference 1)

A typical example of the radial profile of circular velocity of a spiral galaxy is shown in Figure 1. After the initial rise from the disk center, the circular velocity remains essentially constant (or flat). This velocity profile contrasts with that of the solar system, where the orbital velocity of the outer planets about the sun, is markedly slower than that of the inner planets. Various hypotheses have been developed to explain the unexpected galaxy rotation curves. This paper seeks to demonstrate that the gravitational field in the disk plane, is modified in a way which leads to the observed rotation curves of spiral galaxies. The distribution of mass throughout the disk, and the thin, flat disk are the key features which modify the gravitational field. For star systems with a *spherically symmetric* mass distribution, there is a theorem by Newton (Reference 2, p.34) stating that the gravitational force at a given radius is determined solely by the mass *inward* from that radius. The images of spiral galaxies show them to be disks of visible star mass that are clearly *not* spherically symmetric distributions, so the theorem is inapplicable. The model developed in this paper shows that the gravitational force in the disk plane at any radius, must recognize the mass both *inward* and *outward* from that radius.

To construct a model for numerical analysis of the galactic disk, the basic element is a *ring* of matter. Later, the *disk* will be modeled by a series of concentric, coplanar ring elements.

## Gravitational Force in the Ring Plane

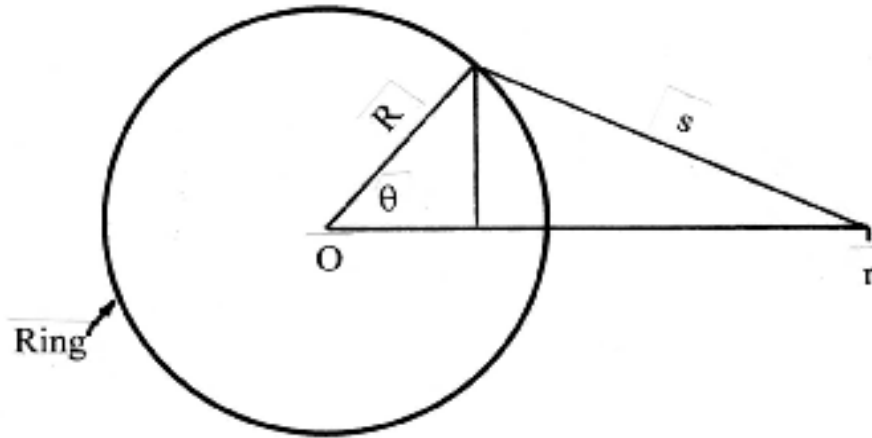


Figure 2. Force Diagram in Ring Plane

If  $F(r)$  is the gravitational force in the plane of a ring element of mass  $\Delta M$ , then

$$F(r) = \frac{G \Delta M}{2\pi} \int_0^{2\pi} \left[ \frac{r - R \cos \theta}{s^3} \right] d\theta \quad (1)$$

$$F(r) = \frac{G \Delta M}{r^2} H(\zeta) \quad \text{where } \zeta = \frac{R}{r} \quad (2)$$

$$H(\zeta) = \frac{1}{2\pi} \int_0^{2\pi} \left[ \frac{(1 - \zeta \cos\theta)}{(1 + \zeta^2 - 2\zeta \cos\theta)^{3/2}} \right] d\theta \quad (3)$$

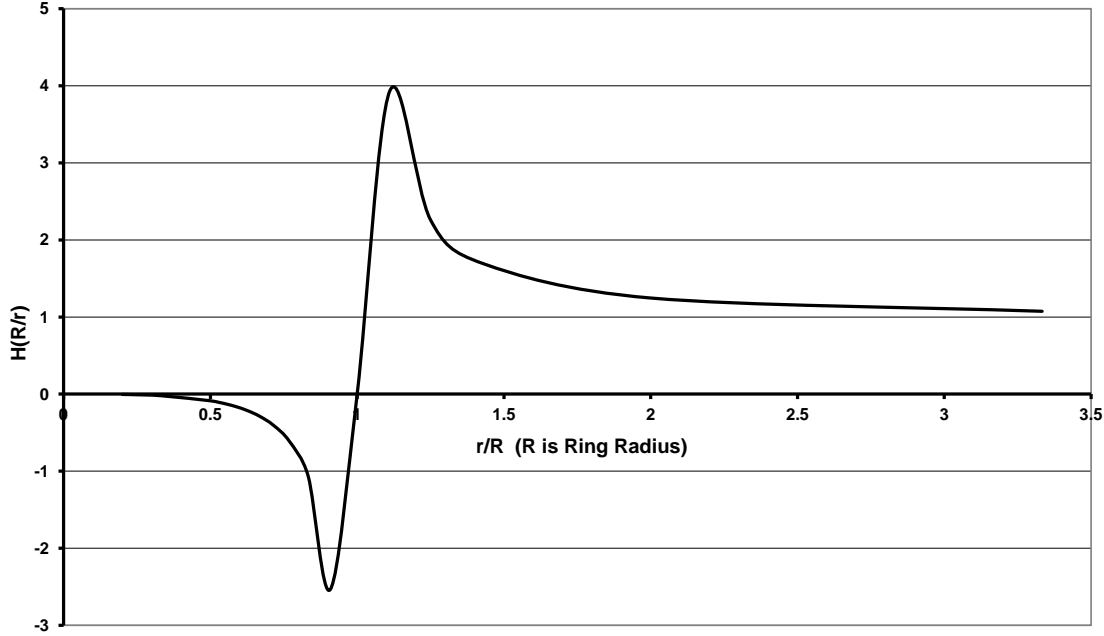


Figure 3.  $H(R/r)$  in Ring Plane

The function  $H(R/r)$  can be integrated numerically to give the result shown in Figure 3. *Beyond* the ring, the force is positive (attracting toward the ring center). It peaks and then subsides, so that from about  $r/R = 2$ , the force will start to match the characteristics of a point mass at the ring center. *Within* the ring, the force is negative (attracting toward the ring, but *away* from the center). There is a negative peak near the ring and then the force subsides to zero at the ring center. (A fuller derivation is provided in Reference 3.)

## Gravitational Forces in the Disk Plane

A model of a thin, axially symmetric disk is constructed with a series of  $n$  concentric, coplanar rings. The basic model comprises a series of 10 equally-spaced rings at  $r/R_m = 0.05, 0.15, 0.25, \dots, \text{etc.} \dots 0.85, 0.95$ . The numerical integration proceeds with summation steps at  $r/R_m = 0.1, 0.2, 0.3, \dots, \text{etc.} \dots 0.9, 1.0$  and further, as desired. Here  $R_m$  denotes the edge or “rim” radius of the disk selected for evaluation. Each ring is allotted its respective fraction of the total mass of the disk within  $R_m$ , in accordance with an assumed mass distribution.

The gravitational force in the disk plane and toward the disk center, at radius  $(r_i/R_m)$  is given by

$$F(r_i/R_m) = \frac{1}{(r_i/R_m)^2} \sum_{i=1}^n \Delta M_i [H(R_m/r_i)] \quad (4)$$

To demonstrate the model, a radial profile of mass distribution in the disk is chosen as a linear variation, such that the slope of the line is -80 percent. This is noted as an “L80” mass distribution and is illustrated in Figure 4.

(Various mass distributions, both linear and non-linear, have been tested to review their effect on the radial velocity profile of the model. Appendix 1 of this paper describes how a specific mass distribution can be derived, which matches the observed rotation curve for a given spiral galaxy.)

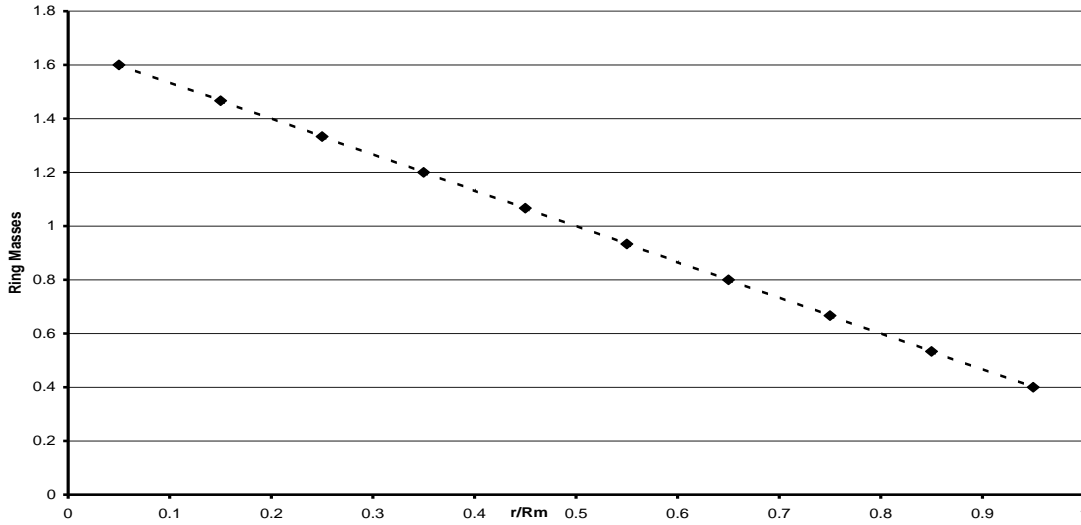


Figure 4. Ring Masses for L80 Mass Distribution

The area of the disk increases with the radius, so the surface mass density at the disk center is several hundred times the value at the disk rim, as illustrated in Figure 5.

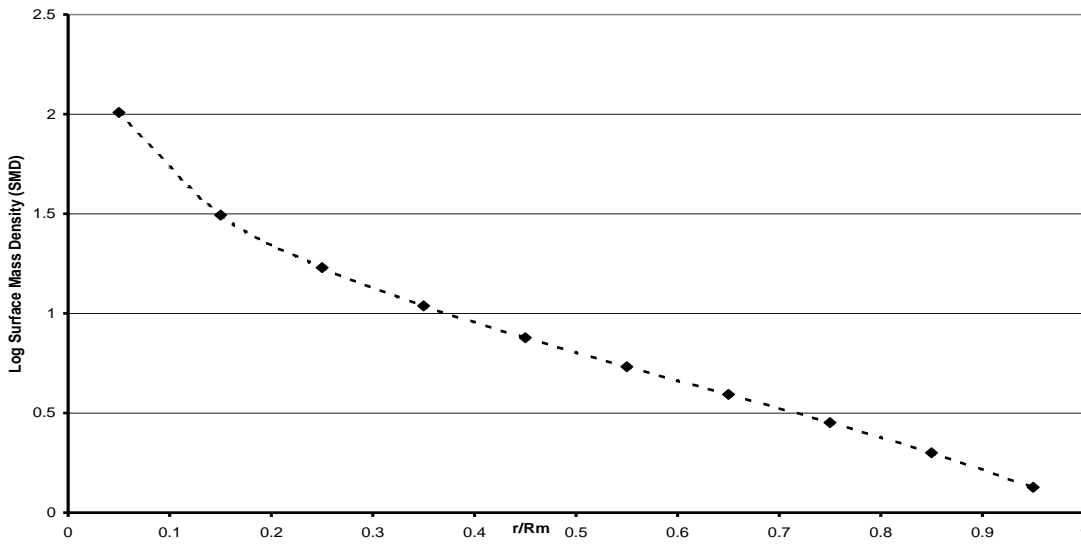


Figure 5. Surface Mass Density for L80 Mass Distribution

The calculated gravitational force in the plane of the disk is shown in Figure 6.

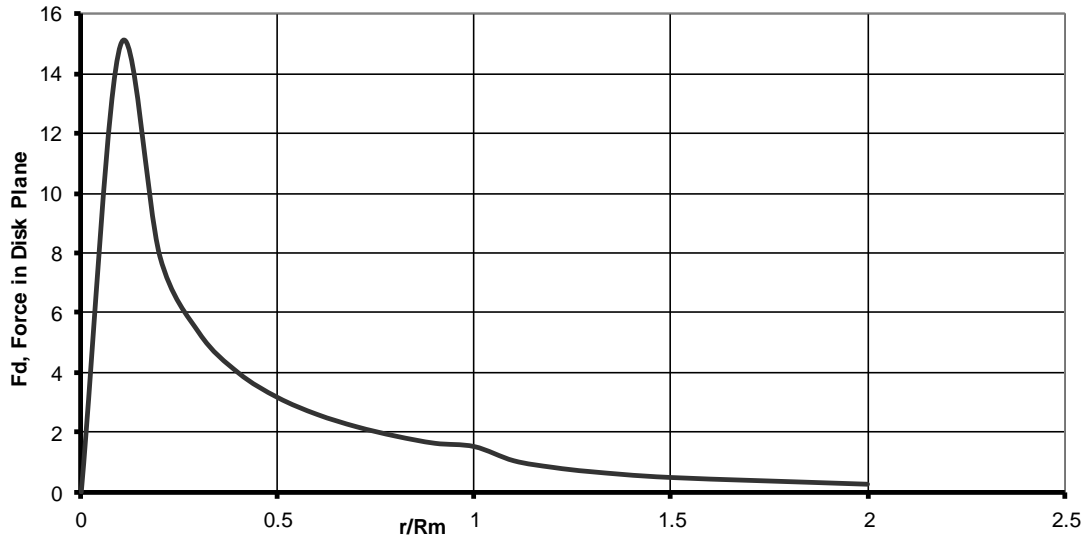


Figure 6. Gravitational Force in Disk Plane for L80 Mass Distribution

Comparing the disk force ( $F_d$ ) to the force from a point mass ( $F_p$ ) located at the disk center, is more instructive, as illustrated in Figure 7.

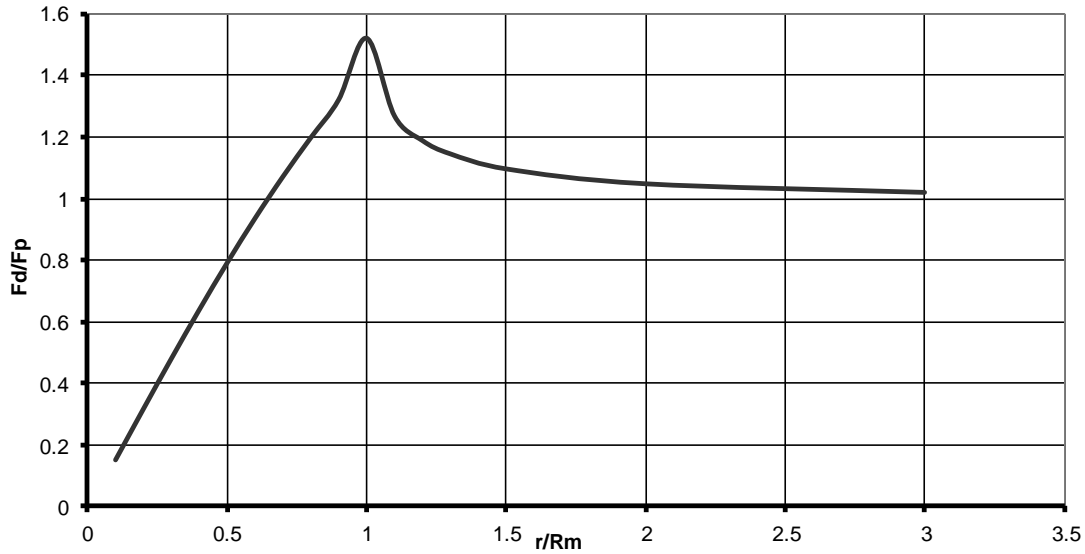


Figure 7. Disk Force vs. Force for Point Mass (L80 Mass Distribution)

For the L80 mass distribution, the disk force factor ( $F_d/F_p$ ) increases from the disk center, passes through unity at about  $r/R_m=0.7$ , and crests to a value of about 1.5 at the disk rim. (The

factor ( $F_d/F_p$ ) peaks nearer the disk center for rotation curves that decline substantially before  $R_m$ ). The enhancement of the gravitational force in the disk plane subsides only gradually beyond  $R_m$ . From about  $r/R_m = 2$ , the ratio ( $F_d/F_p$ ) approaches unity, and the gravitational force then proceeds as a Keplerian decline.

### Circular Velocities in the Disk

The circular velocity  $V_i$  at radius  $\{r_i/R_m\}$  is given by

$$V_i^2 = \{F(r_i/R_m)\}\{r_i/R_m\} \tag{5}$$

Balancing the centrifugal and gravitational forces at any radius produces the rotation curve shown in Figure 8. This relatively flat profile corresponds to the observed velocity profiles for many spiral galaxies. Note that in the normalized rotation curve for the L80 mass distribution, the circular velocity at  $r/R_m = 1$  is  $V_m^* = 1.24$

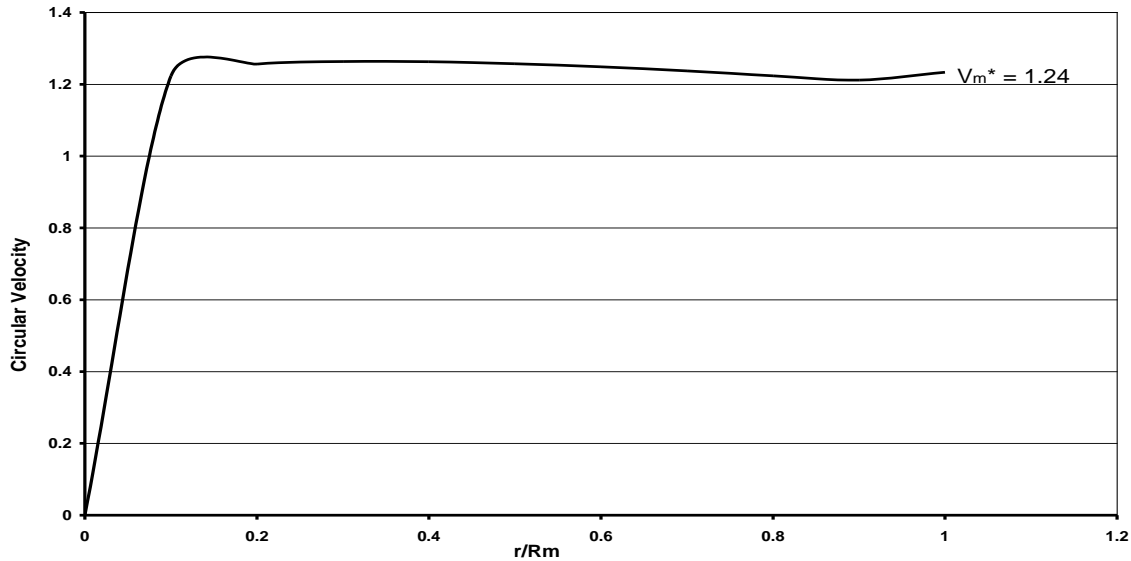


Figure 8. Rotation Curve for L80 Mass Distribution

### Estimation of the Mass of a Spiral Galaxy

The total disk galaxy mass includes both the optically observable star matter and the (radio-observable) gas clouds, out to the farthest measured data point. With this simple model, no attempt is made to take specific account of features such as the nuclear bulge, possible black hole, etc. at the center of a typical spiral galaxy. However, with the presumption that the rotation curve is solely dependent on the mass distribution in the disk, the derived mass distribution may nevertheless have given some recognition to these other features.

The enclosed galaxy mass within the radial limit of  $R_m$ , is given by

$$M_{R_m} = \frac{(R_m \times V_m^2)}{\{G \times (F_d/F_p)_m\}} \quad (6)$$

For equation 6 to be valid,  $R_m$  must be located at the *outermost* radial point of velocity data measurement (and where disk mass density has effectively declined to zero). The gravitational constant for the purposes of this paper is given by

$$G \approx 4.3 \times 10^{-6} \text{ kpc (km/s)}^2 M_{\odot}^{-1}$$

Using a “trial and error” adjustment of the model ring masses has proven capable of achieving a reasonable match with real spiral galaxy velocity profiles. Two examples (NGC 3198 and M31, Andromeda) are provided below. (A detailed description of the convergence method used to determine the individual ring masses that produce a rotation curve matching a given spiral galaxy (NGC 4736), is provided in Appendix 1 of this paper.)

## Model for NGC 3198

The rotation curve for NGC 3198 (as taken from Reference 2, p.601) is shown in Figure 9.

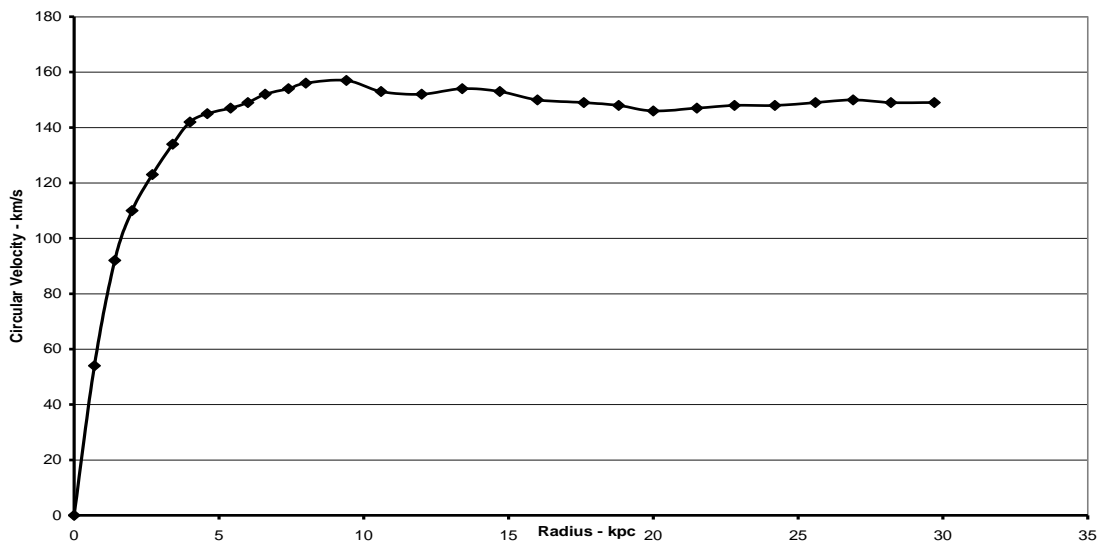


Figure 9. Radial Profile of Circular Velocity for NGC 3198

The basic Model which utilized 10 rings, has been increased to 12 rings, which gives improved modeling capability near the disk center. The matching of the velocity curves for NGC 3198 and the Model is shown in Figure 10.

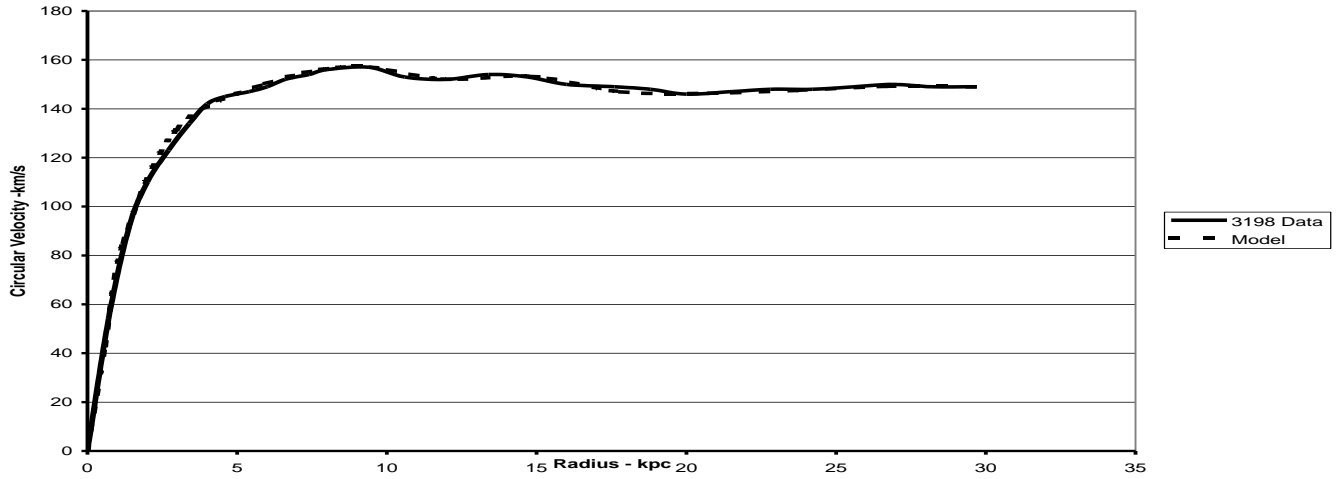


Figure 10. Velocity Profiles of NGC 3198 and Model

The relevant data for estimation of enclosed galaxy mass within  $R_m$  are:

$$R_m = 29.7 \text{ kpc} \quad V_m = 149 \text{ km/s} \quad (F_d/F_p)_m = 1.527$$

$$\text{Mass of NGC 3198 is.....} \quad M_m = 10.04 \times 10^{10} M_\odot$$

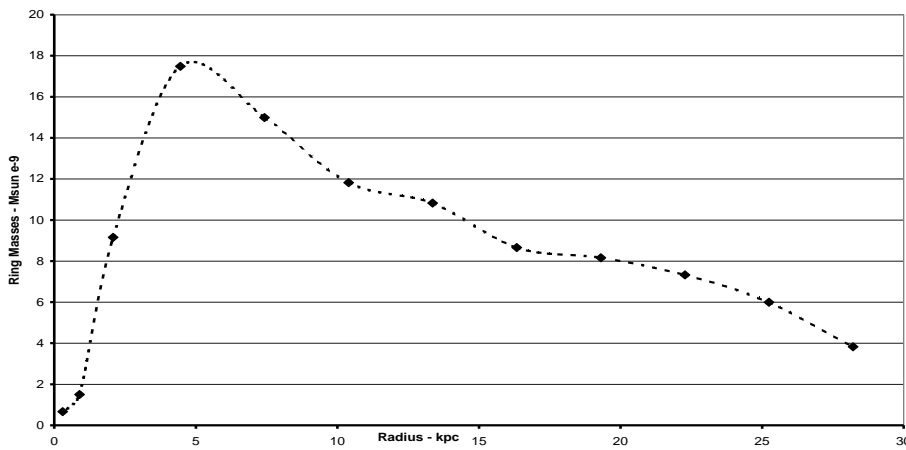


Figure 11. Ring Mass Distribution for NGC 3198 Model



Figure 11 gives the ring mass distribution that produced the model velocity profile (shown as the dashed line in Figure 10) .

Figure 12 translates the ring mass distribution of Figure 11 into a Surface Mass Density format. Figure 13 illustrates the buildup of enclosed galaxy mass for the NGC 3198 Model.

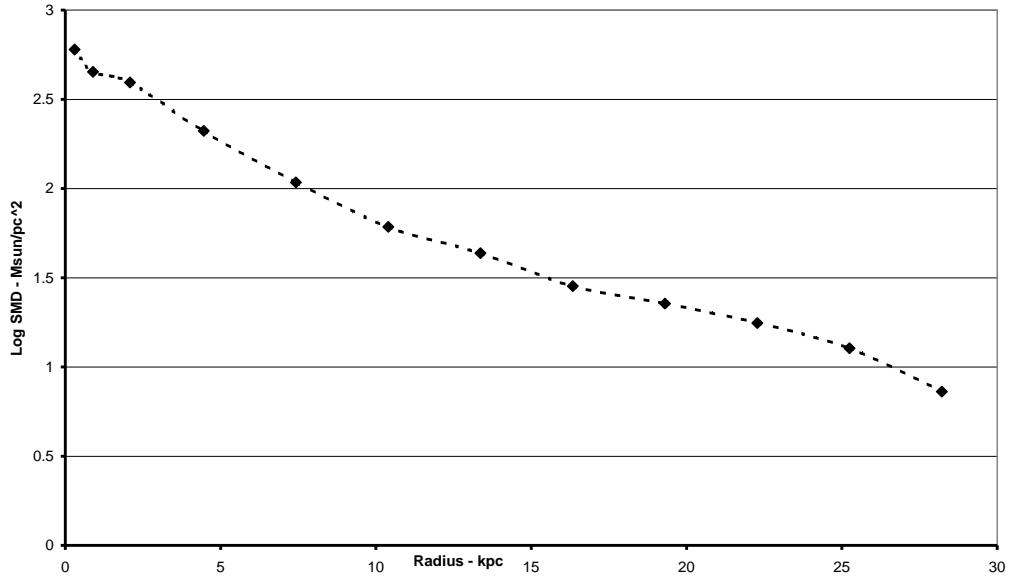


Figure 12. Surface Mass Density for NGC 3198 Model

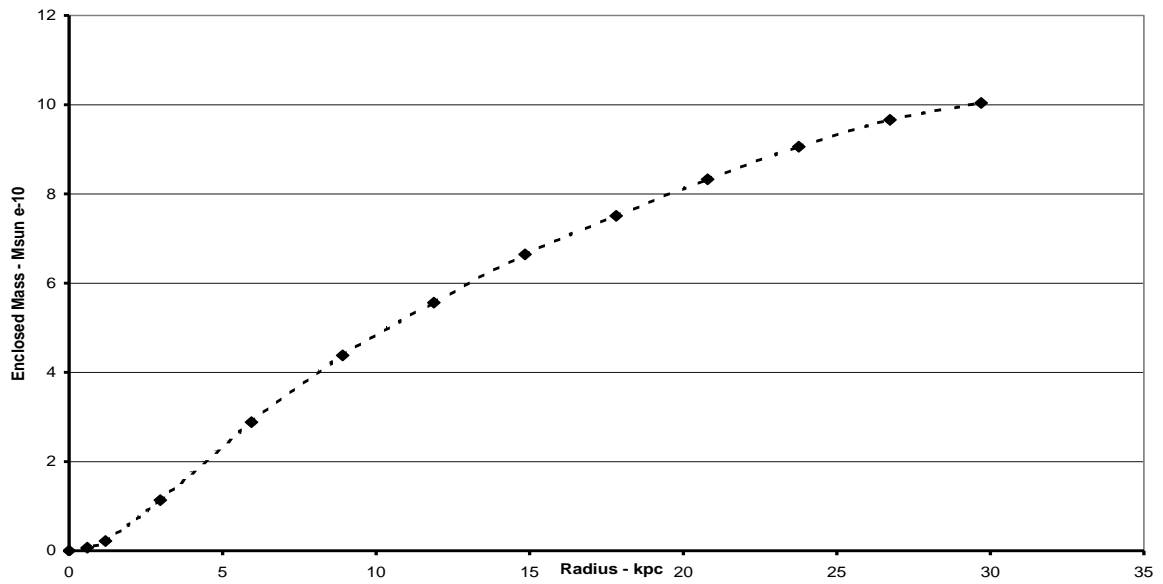


Figure 13. Enclosed Mass for NGC 3198 Model

## Model for M31 (Andromeda Galaxy)

The rotation curve values for M31 (as taken from Reference 4.) are shown in Figure 14.

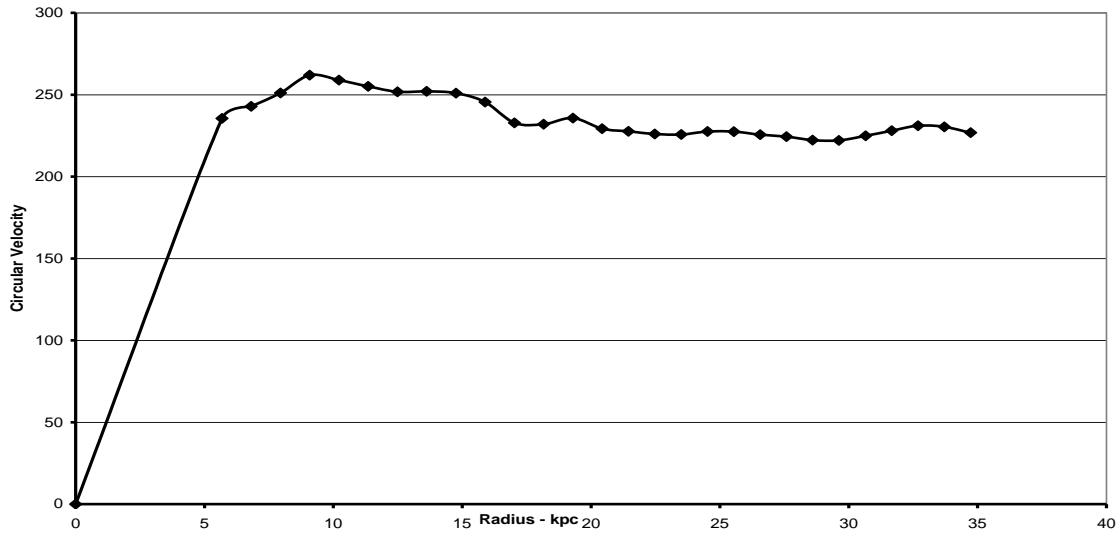


Figure 14. Radial Profile of Circular Velocity for M31

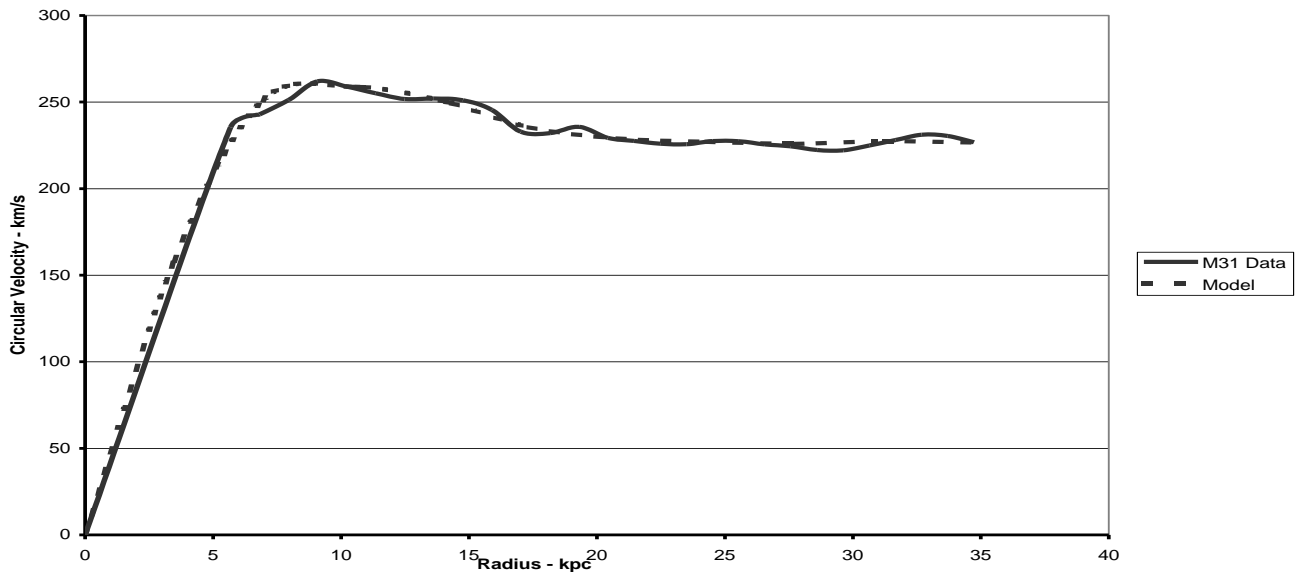


Figure 15. Velocity Profiles of M31 and Model

Figure 15 shows the Model achieving a reasonable match with the rotation curve of M31.

The relevant data for estimating the galaxy mass of M31 enclosed within  $R_m$  are:

$$R_m = 34.73 \text{ kpc} \quad V_m = 226.8 \text{ km/s} \quad (F_d/F_p)_m = 1.485$$

$$\text{Mass of M31 is....} \quad M_m = 28.0 \times 10^{10} M_\odot$$

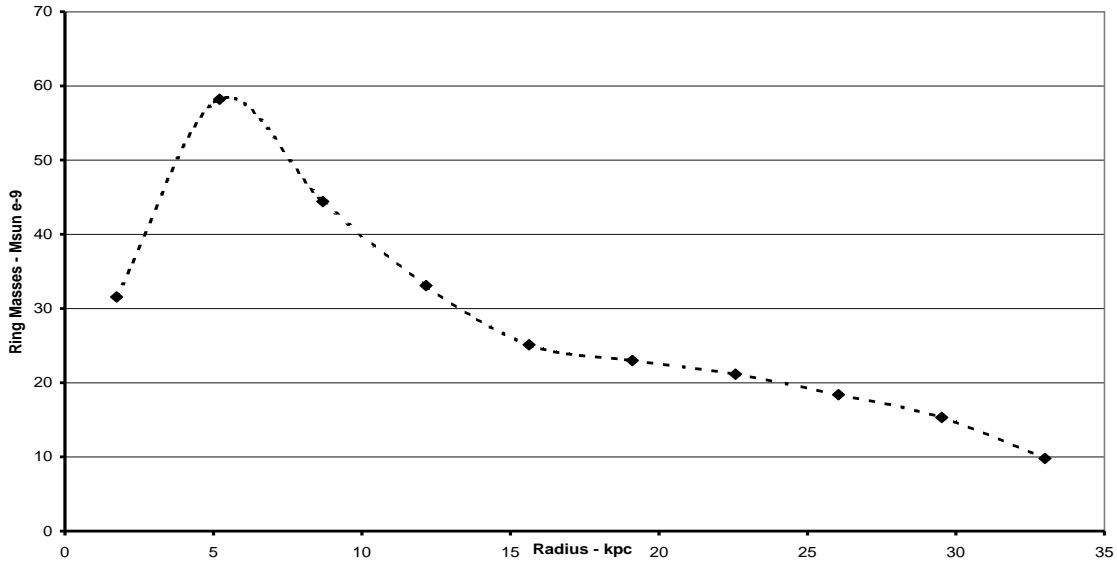


Figure 16. Ring Mass Distribution for M31 Model

Figure 16 shows the ring mass distribution that produced the model velocity profile (shown as the dashed line in Figure 15). Figure 17 translates the ring mass distribution of Figure 16 into a Surface Mass Density format. Figure 18 illustrates the buildup of Enclosed Galaxy Mass for the M31 Model.

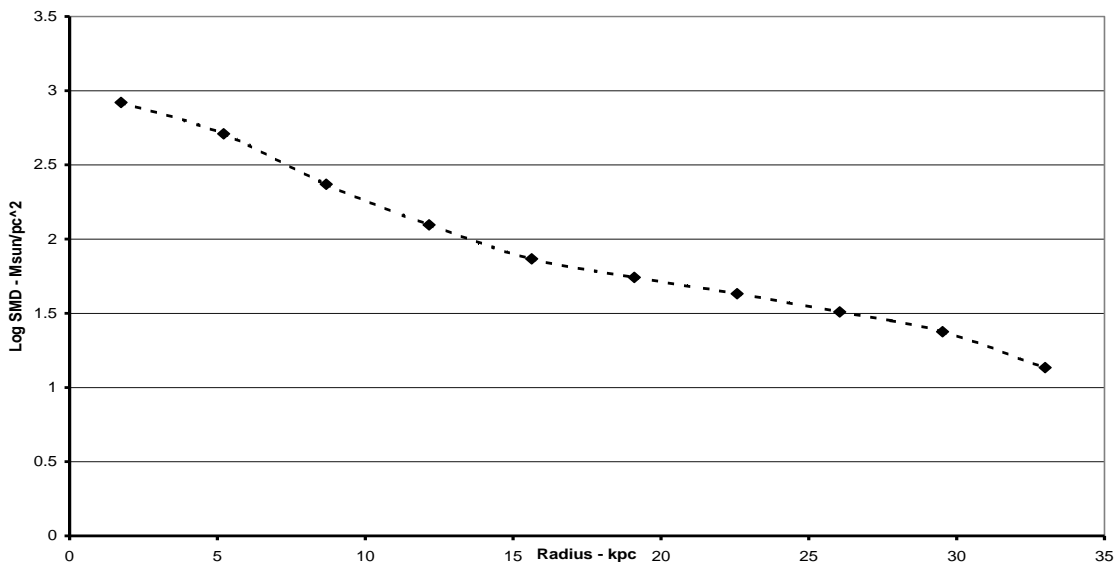


Figure 17. Surface Mass Density for M31 Model

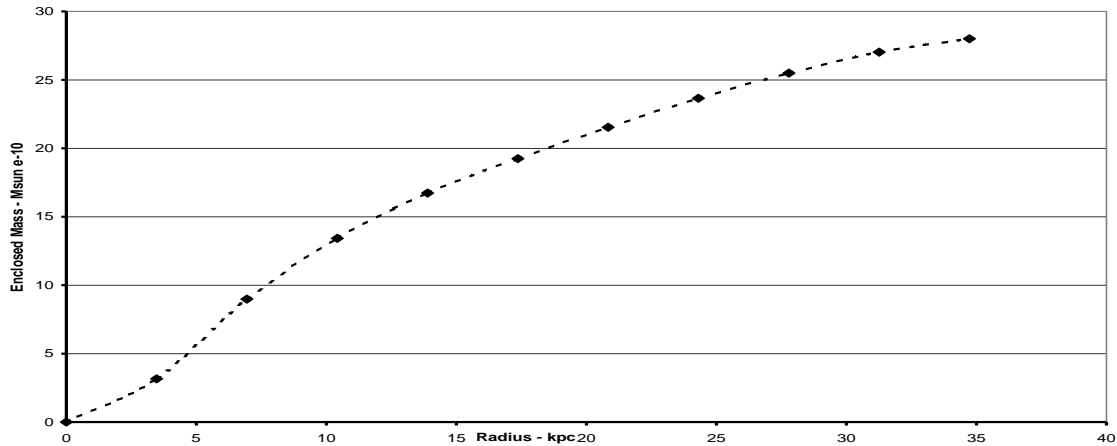


Figure 18. Enclosed Mass for M31 Model

## Discussion

### Dark Matter Models

The “traditional” method of estimating the mass distribution in spiral galaxies has been to decompose the galaxy into its various features such as disk, gas, dark matter halo, etc. as indicated in Figure 19.

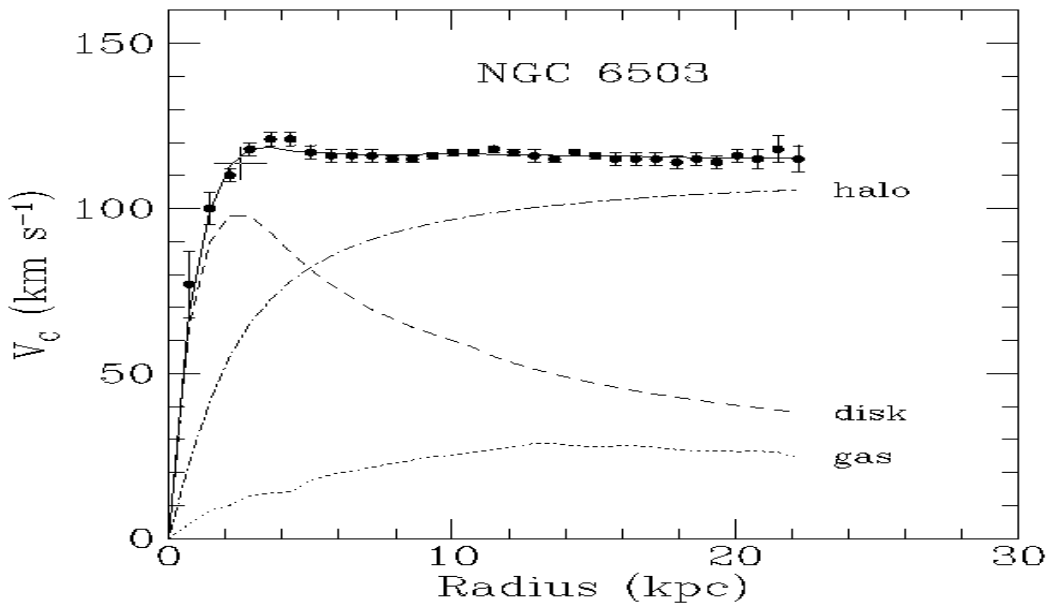


Figure 19. Typical Mass Components of (traditional) Galaxy Model

Allowing the various components to separately satisfy certain assumptions (such as mass-luminosity relationships) produced estimates of individual component mass and its radial

distribution. When these contributions were summed, it became apparent that there was insufficient observable component mass to produce the flat rotation curve. The halo of dark matter was introduced to make up for this apparent deficiency in observable mass in the outer region of the spiral galaxy model. The proportion of dark matter mass to estimated total galaxy mass could be as high as 50 percent (or more) in these traditional studies.

## Newtonian Model

In contrast to the techniques and assumptions adopted in the traditional studies, the proposed Newtonian model exclusively uses the observed rotation curves to derive the radial profiles of mass distribution in the disk. There have been a number of other models (a list is provided in Reference 7.) that also are based exclusively on the observed rotation curves of spiral galaxies. These models do not utilize the measured light distribution, and so have no need to make assumptions about the possible coupling of mass surface density and the measured surface brightness profile.

Spiral Galaxies have substantial thickness in their disks, as is apparent in any edge-on images. So the thin, flattened disk model (of 10 or 12 rings) proposed in this paper may be considered simplistic. However, various approximations are utilized in measuring the rotation curve data for any given spiral galaxy, so the current model is probably adequate for predicting the enclosed galaxy mass and the mass distribution in the disk. Appendix 2 provides summary details of several spiral galaxies that have been analyzed using the Newtonian model proposed in this paper. Table 1 of Appendix 2 shows that the force factor  $(F_d/F_p)_m$  and the average surface mass density  $(SMD_{av})$  vary over a considerable range, depending on the rotation curve of the individual spiral galaxy.

The predictions of galaxy mass distribution may provide an approach to validating the model. The force field above and below the galaxy disk (within a spherical volume extending to a radius of say,  $R_m$ ) should be anisotropic, in conformity with the predicted surface mass density profile of the disk. This may be confirmable by examining the orbits of such tracer objects as globular clusters, planetary nebulae, etc.

## Acknowledgments

The author gratefully acknowledges the many insights and direction provided him by the listed references.

## References

1. *Rotation Curves of Spiral Galaxies* Sofue, Y. and Rubin, V. 2001, Ann. Rev. Astron. Astrophys. 39,137  
<http://www.ioa.s.u-tokyo.ac.jp/~sofue/h-rot.htm>
2. *Galactic Dynamics* Binney, J. and Tremain, S. 1987  
Princeton University Press

3. *Gravitational Force in a Thin Disk of Matter* Thayer Watkins,  
San Jose State University, California  
<http://www.applet-magic.com/gravdisk.htm>
4. *Extended HI Rotation Curve and Mass Distribution of M31*  
Claude Carignan, Laurent Chemin, Walter K. Huchtmeier, and  
Felix J. Lockman 6 Mar 2006  
[arXiv:astro-ph/0603143v1](http://arxiv.org/abs/astro-ph/0603143v1)
5. *Is dark matter present in NGC 4736? An iterative spectral method of finding mass  
distribution in spiral galaxies* Joanna Jalocho, Lukasz Bratek, Marek Kutschera  
The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of  
Science, Poland. 30 Jan 2008  
[arXiv:astro-ph/0611113v3](http://arxiv.org/abs/astro-ph/0611113v3)
6. *The Extended Rotation Curve and the Dark Matter Halo of M33*  
Edvige Corbelli, Paolo Salucci, Sep 1999  
[arXiv:astro-ph/9909252v1](http://arxiv.org/abs/astro-ph/9909252v1)
7. *Galactic Rotation Described with Various Thin-Disk Gravitational Models*  
James Q. Feng and C.F. Gallo, 1 April 2008  
[arXiv:0804.0217v1](http://arxiv.org/abs/0804.0217v1) [astro-ph.GA]
8. *The mass distribution in early type disk galaxies*  
Edo Noordermeer, Thijs van der Hulst, Oct 2003  
[arXiv:astro-ph/0310868v1](http://arxiv.org/abs/astro-ph/0310868v1)
9. *HI Study of the NGC 6744 System*  
Stuart D. Ryder, Wilfred Walsh and David Malin  
Publ. Astron. Soc. Aust., 1999, 16, 84-8
10. *The published extended rotation curves of spiral galaxies: etc.*  
R.H. Sanders 1996  
[arXiv:astro-ph/9606089v1](http://arxiv.org/abs/astro-ph/9606089v1)
11. *Galaxy Rotation Curves without Non-Baryonic Dark Matter*  
J.R. Brownstein and J.W. Moffat, Sep 2005  
[arXiv:astro-ph/0506370v4](http://arxiv.org/abs/astro-ph/0506370v4)
12. *General Relativity Resolves Galactic Rotation without Exotic Dark Matter*  
F.I. Cooperstock and S.Tieu, Jul 2005  
[arXiv:astro-ph/0507619v1](http://arxiv.org/abs/astro-ph/0507619v1)
13. *Spiral Galaxy Rotation Curves determined from Carmelian General Relativity*  
John G. Hartnett, Feb 2008  
[arXiv:astro-ph/0511756v3](http://arxiv.org/abs/astro-ph/0511756v3)

## Appendix 1.

This Appendix describes the convergent method used to derive the discrete ring mass distribution that produces a model velocity profile that satisfactorily matches the rotation curve for a given spiral galaxy. The rotation curve for NGC 4736 (from Reference 5) is used to demonstrate the method and is illustrated in Figure 20.

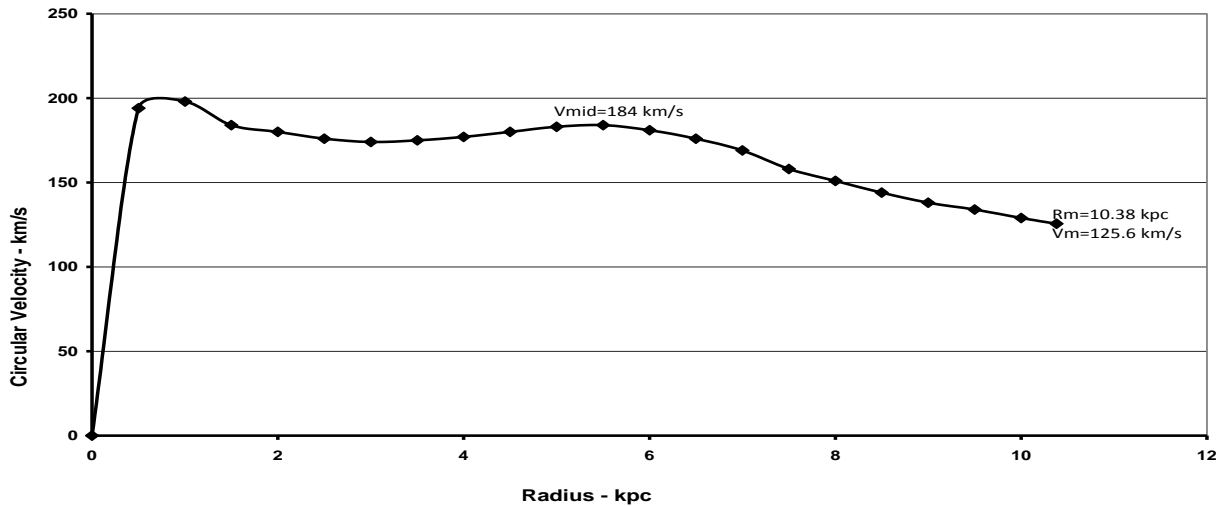


Figure 20. Radial Profile of Circular Velocity for NGC 4736

To construct a “target” rotation curve for NGC 4736, the data in Figure 20 are normalized. This means that  $R_m=1$  corresponds to 10.38 kpc. Normalizing the velocity ( $V_m = 125.6$  km/s) needs careful treatment to conform with the model. Figure 8 of this paper shows that for a relatively flat rotation curve, the model  $V_m^*$  equals about 1.24.

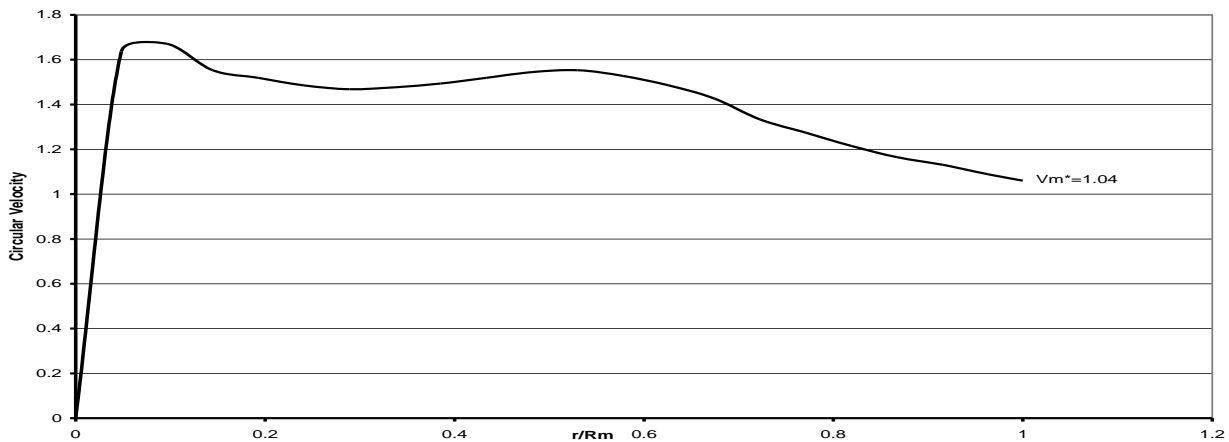


Figure 21. Normalized Velocity Profile for NGC 4736

Experience with various mass distributions has shown the model normalized  $V_m^*$  can be approximated by

$$V_m^* \approx 0.6 (V_m/V_{mid}) + 0.63 \quad (7)$$

where  $V_{\text{mid}}$  is the velocity of the target rotation curve at about  $r/R_m = 0.5$ . For NGC 4736, equation 7 suggests an initial model  $V_m^* \approx 1.04$ , and it helps the convergence process to normalize the target rotation curve to reflect this value, as indicated in Figure 21.

A modified L80 mass distribution (see Figure 4.) was chosen as the *starting* model mass distribution for this demonstration. A chart was constructed which compared the velocity (corresponding to the model mass distribution) with the target velocity of Figure 21. The velocity chart (Figure 23) immediately shows the effect of any changes to the model ring mass distribution, so the convergent process becomes interactive.

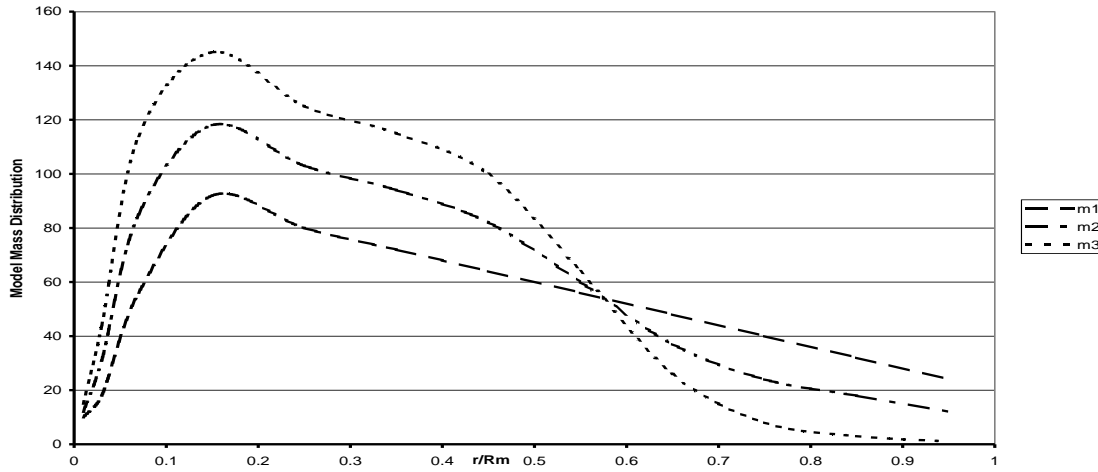


Figure 22. Stages of Model Mass Distribution for NGC 4736

Figure 22 shows the successive stages in the mass distribution adjustment, and Figure 23 shows the corresponding stages in the model velocity profile as it converges with the target profile of NGC 4736. The process consists of changing one or two individual ring mass values and immediately reviewing whether the changes are favorable and leading to convergence.

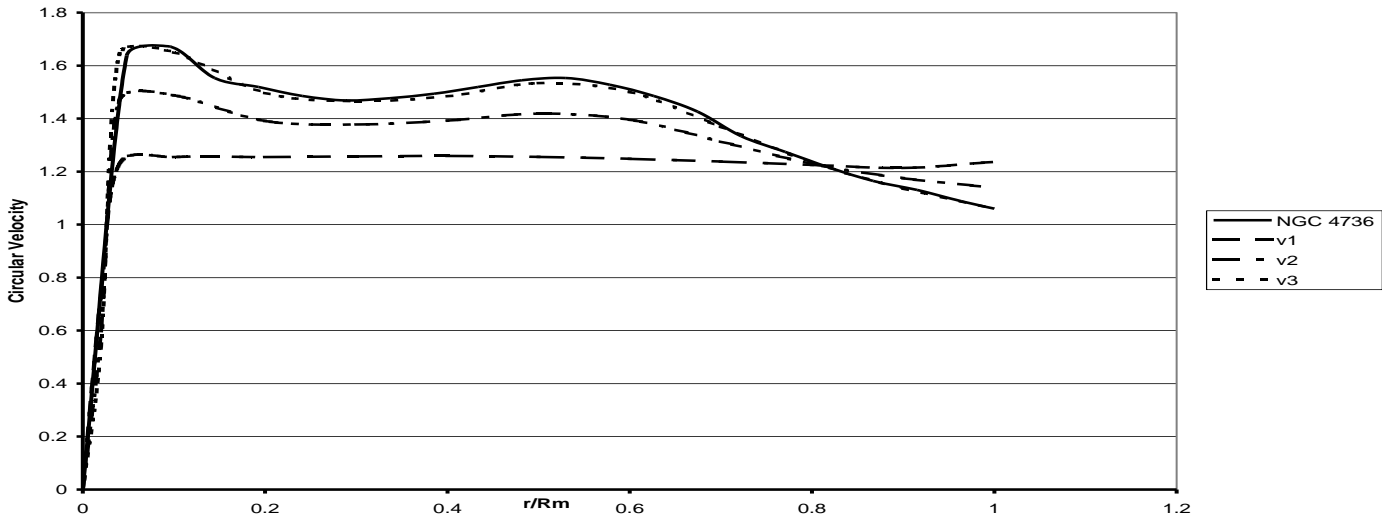


Figure 23. Convergence of Model Velocity to match NGC 4736



Since the model is recognizing mass values both inward and outward from the radial point of interest, the method tends to be counter-intuitive, but becomes manageable with practice. The three stages shown in Figures 22 and 23 involved a total of about 100 individual ring mass adjustments and was completed in about 20 minutes. As the convergent process approaches its final adjustment, the model data indicates that the velocity at  $r/R_m = 1$  is settling at  $V_m^* = 1.06$ , so the target rotation curve is scaled up to reflect this small change. With only 12 discrete mass points, it is obvious that this simplistic match with the target rotation curve will be approximate, but this may be satisfactory for many situations.

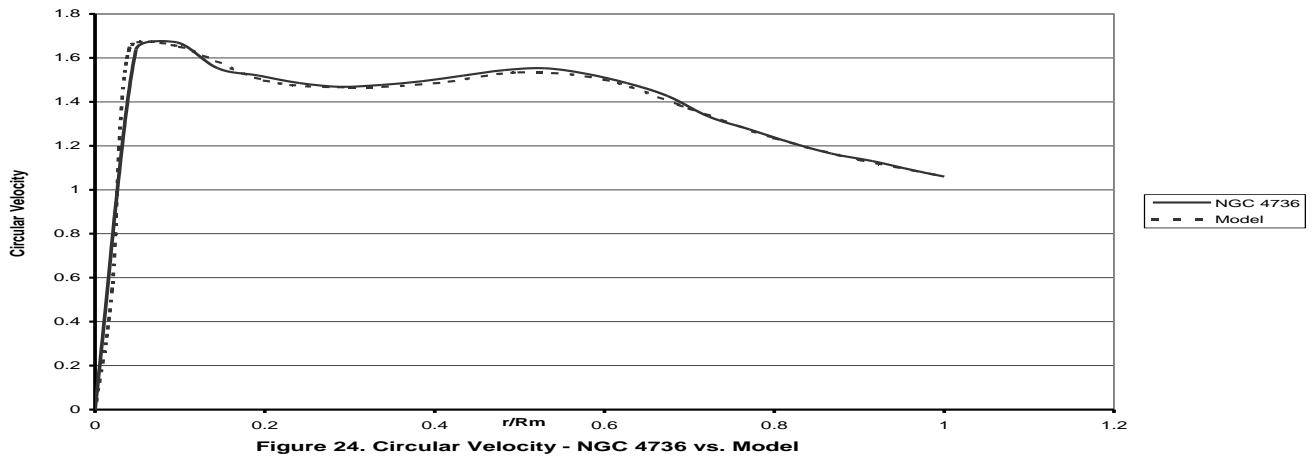
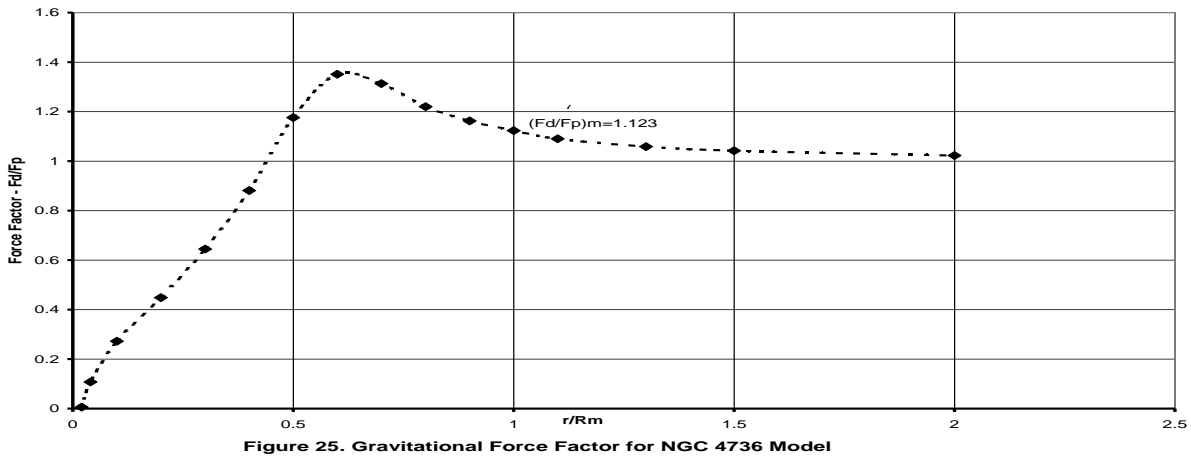


Figure 24 shows the final match between the model velocity profile and NGC 4736 data.



The gravitational force factor ( $F_d/F_p$ ) in the disk is calculated for the final mass distribution, and is shown in Figure 25. The factor crests with a value of about 1.38 at  $r/R_m = 0.6$ , and has a value of 1.123 at  $r/R_m = 1$ .

The relevant data for estimating the galaxy mass of NGC 4736 enclosed within  $R_m$  are:

$$R_m = 10.38 \text{ kpc} \quad V_m = 125.6 \text{ km/s} \quad (F_d/F_p)_m = 1.123$$

$$\text{Mass of NGC 4736 is...} \quad M_m = 3.39 \times 10^{10} M_\odot$$

Figure 26 shows the ring mass distribution that produced the model velocity profile that matched the NGC 4736 rotation curve. Figure 27 translates this ring mass distribution into a Surface Mass Density format. This SMD corresponds closely with that shown in Figure 4 of Reference 5, which was derived by a different methodology, thus providing some validation for the model proposed in this paper.

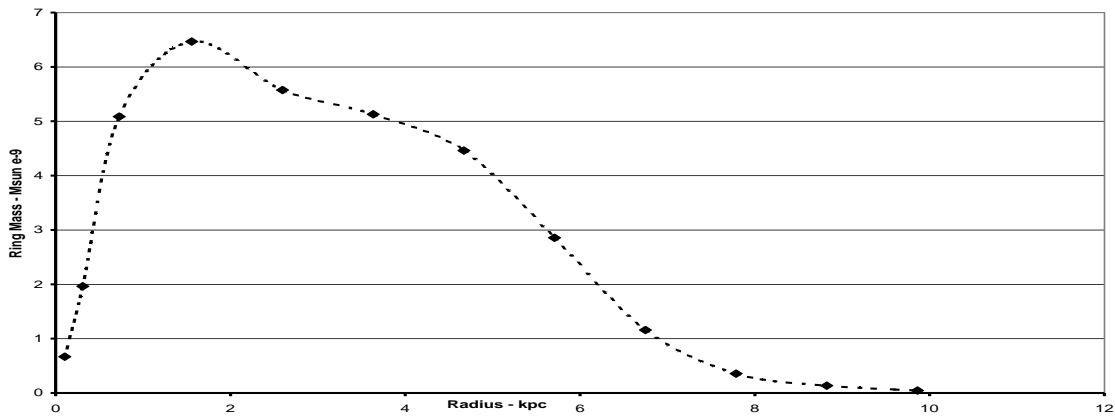


Figure 26. Ring Mass Distribution for NGC 4736 Model

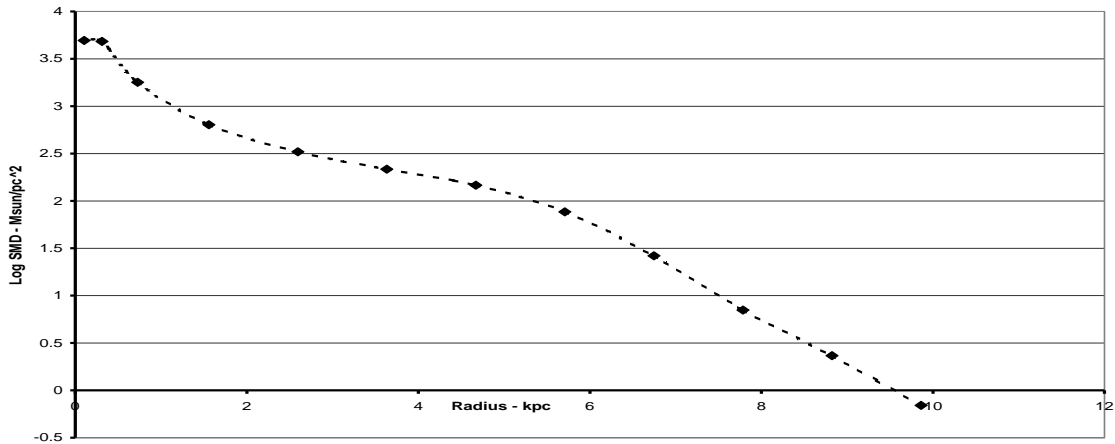


Figure 27. Surface Mass Density for NGC 4736 Model

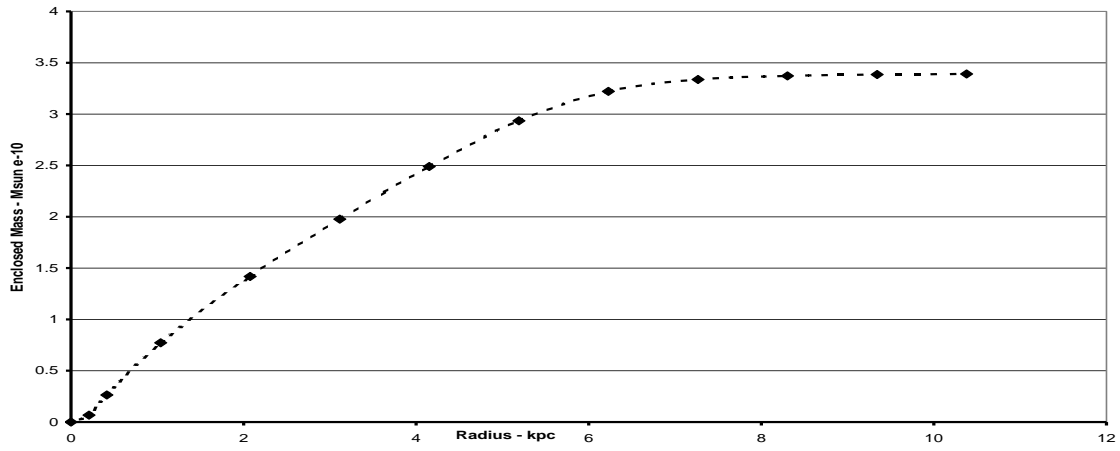


Figure 28 shows the buildup of the enclosed galaxy mass for the NGC 4736 Model. The enclosed galaxy mass estimate of  $3.39 \times 10^{10} M_{\odot}$  is close to the value of  $3.43 \times 10^{10} M_{\odot}$  derived by a different methodology in Reference 5.

## Appendix 2. Basic characteristics of several Spiral Galaxies

The mass distributions and enclosed galaxy mass for several spiral galaxies (as noted in Table 1 below) have been analyzed using the Model proposed in this paper. In the following pages, three basic charts are provided for each galaxy:

- A chart showing the published rotation curve data for the galaxy (solid line) and the model matching curve (dashed line).
- A chart showing the associated ring mass distribution predicted by the model.
- A chart showing the associated (log) surface mass density distribution.

The objective of this appendix is to demonstrate the capability of the Model in analyzing a selection of spiral galaxies with widely varying rotation curve shapes. Table 1 provides the summary details of the spiral galaxies that have been analyzed in this paper and the appendices.

Galaxy	Type	$R_m$ kpc	$V_m$ km/s	$M_m$ $10^{10}M_\odot$	$SMD_{av}$ $M_\odot/pc^2$	$(F_d/F_p)_m$
Milky Way <sup>12</sup>	Sb	29.3	212.0	20.9	77.5	1.46
M31 <sup>4</sup>	Sb	34.73	226.8	28.0	73.9	1.48
NGC 598 <sup>6</sup>	Scd	15.4	136.0	3.72	49.9	1.78
NGC 2590 <sup>13</sup>	Sbc	15.8	275.0	17.7	226.0	1.57
NGC 3031 <sup>12</sup>	Sab	21.0	169.5	10.8	78.0	1.29
NGC 3198 <sup>2</sup>	Sc	29.7	149.0	10.0	36.2	1.53
NGC 4157 <sup>11</sup>	Sb	30.8	185	16.2	54.4	1.51
NGC 4736 <sup>5</sup>	Sab	10.38	125.6	3.39	100.1	1.12
NGC 5533 <sup>8</sup>	Sab	97.5	230.0	80.0	26.8	1.50
NGC 6744 <sup>9</sup>	Sbc	39.2	217.0	25.1	52.0	1.71
NGC 6946 <sup>10</sup>	Scd	29.4	159.0	11.9	43.8	1.45
UGC 2885 <sup>10</sup>	Sbc	72.6	298.0	95.6	57.7	1.57

The superscripts next to the galaxy designation indicate the Reference which provided the published rotation curve data. The Average Surface Mass Density ( $SMD_{av}$ ) is the enclosed galaxy mass ( $M_m$ ) divided by the circular disk area bounded by  $R_m$ .

**Table 1. Basic characteristics of several Spiral Galaxies**

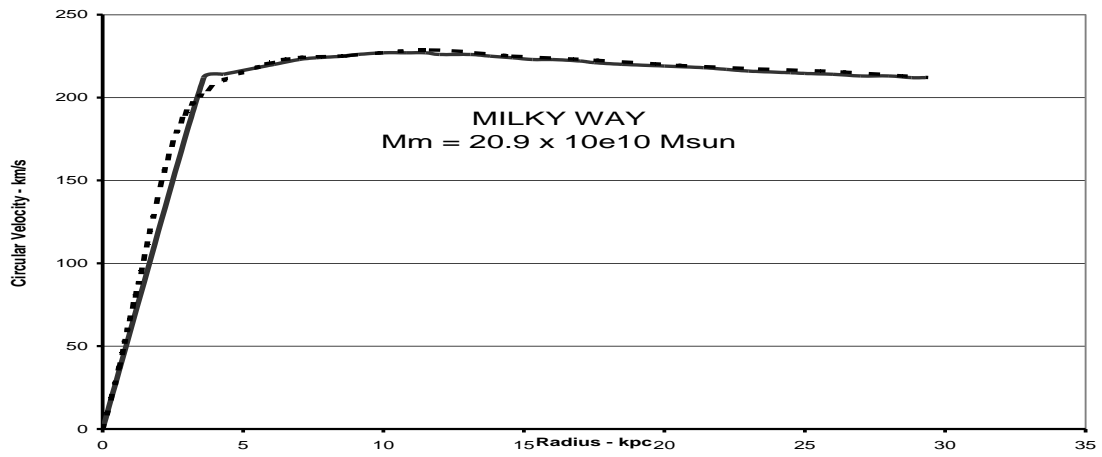


Figure 29a. Velocity Profiles of Milky Way and Model

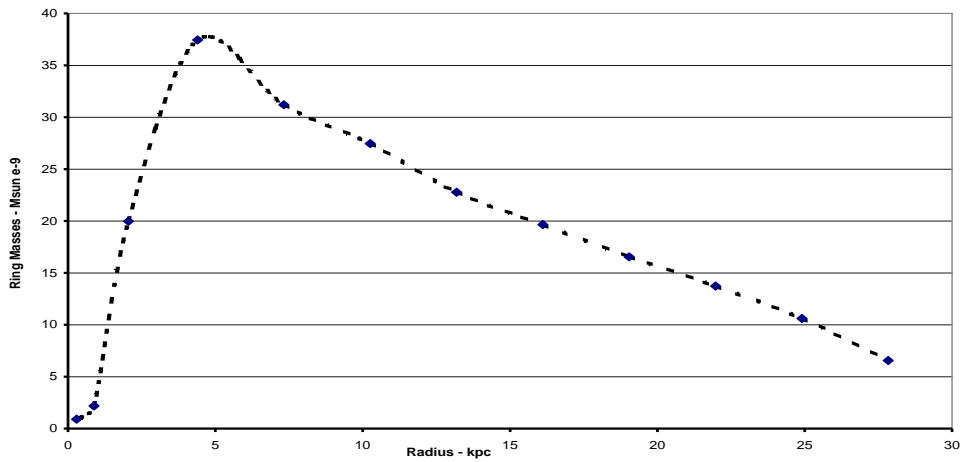


Figure 29b. Ring Mass Distribution for Milky Way Model

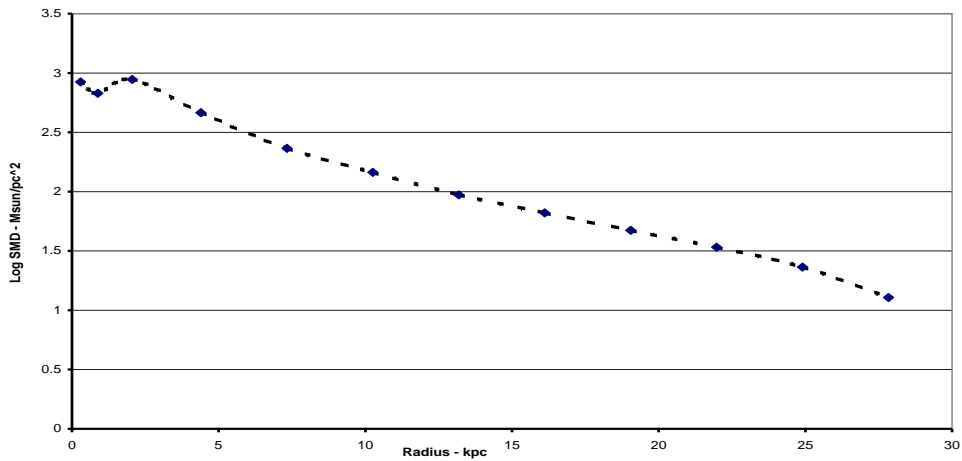


Figure 29c. Surface Mass Density for Milky Way Model

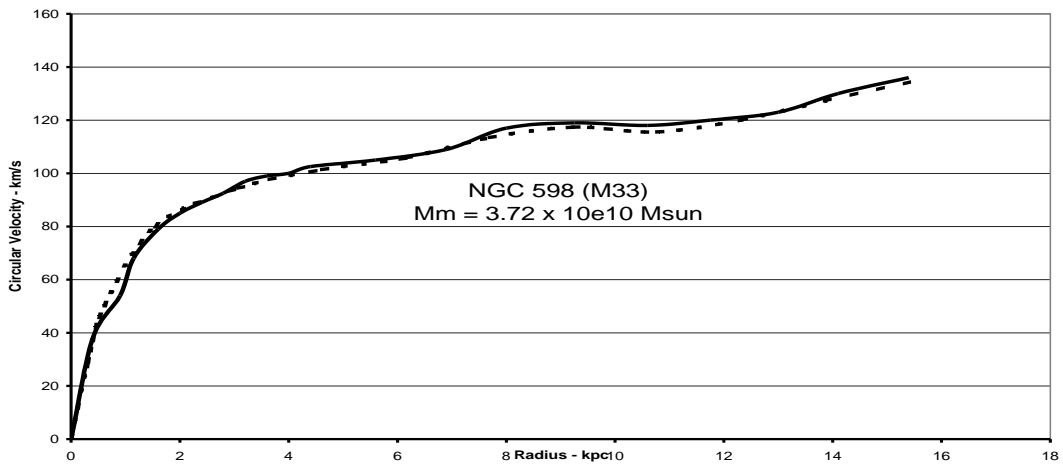


Figure 30a. Velocity Profiles for M33 and Model

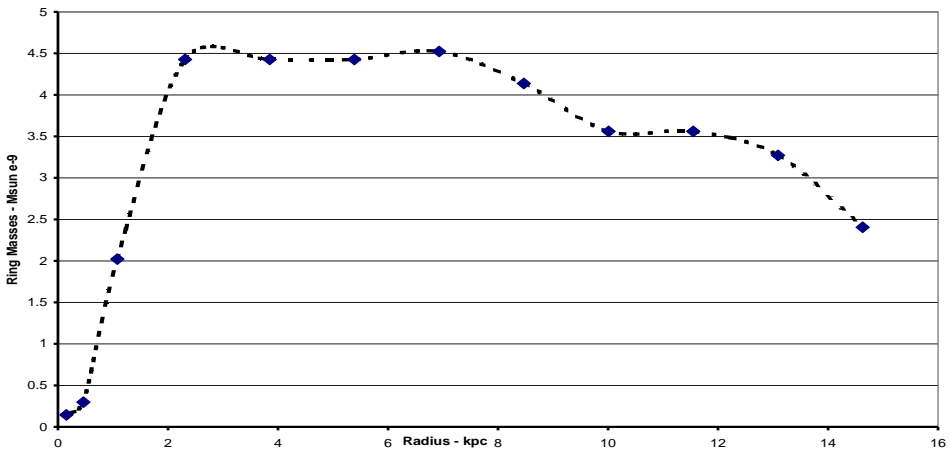


Figure 30b. Ring Mass Distribution for M33 Model

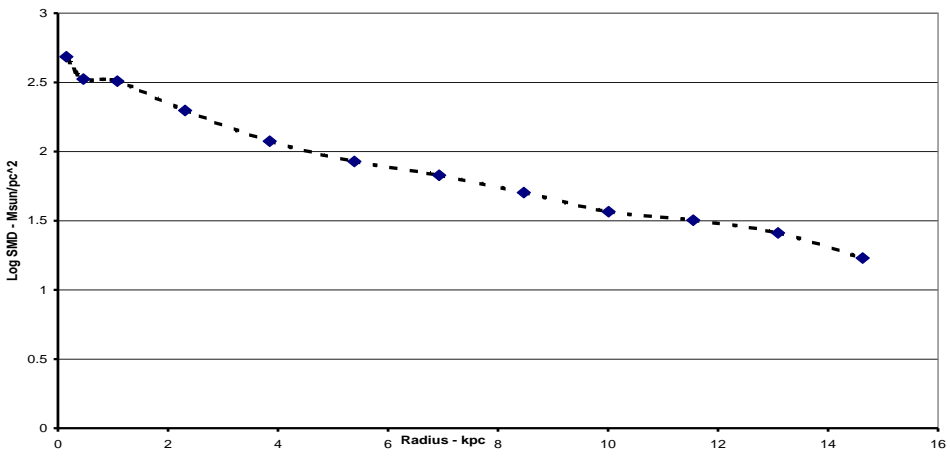
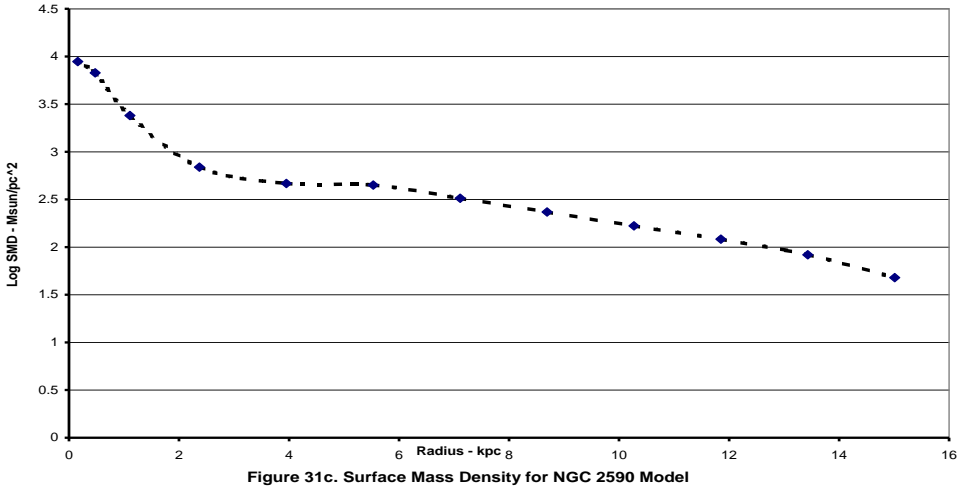
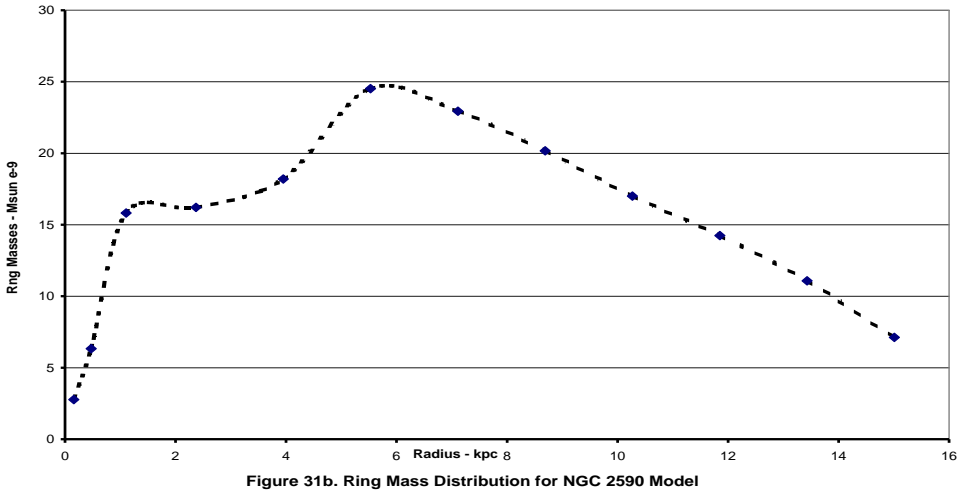
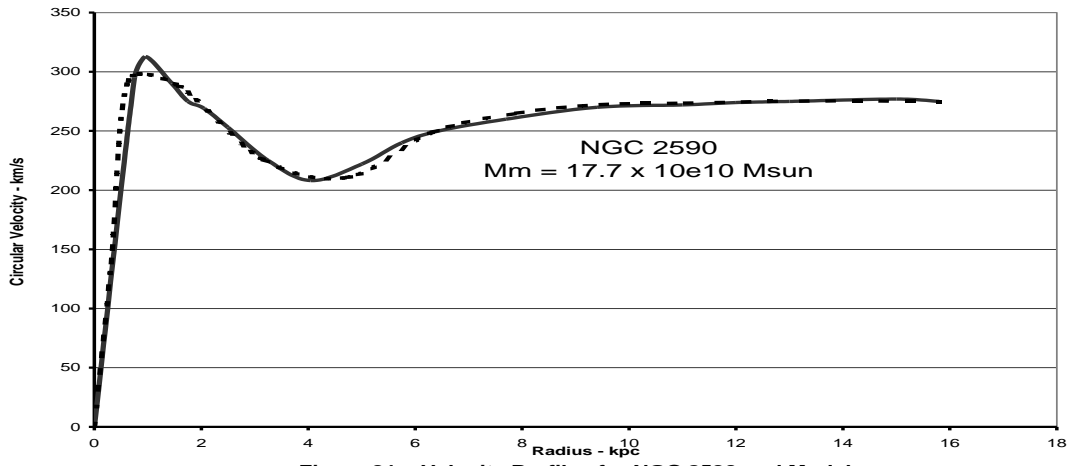


Figure 30c. Surface Mass Density for M33 Model



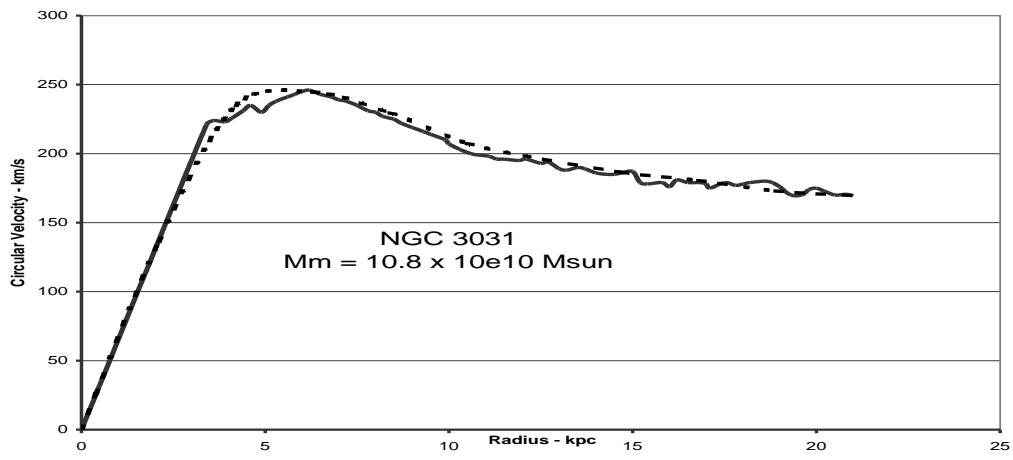


Figure 32a. Velocity Profiles of NGC 3031 and Model

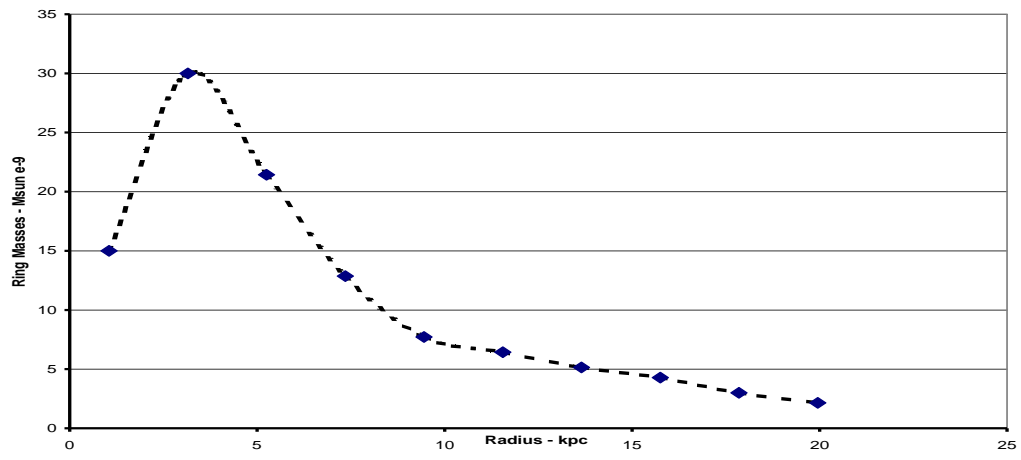


Figure 32b. Ring Mass Distribution for NGC 3031 Model

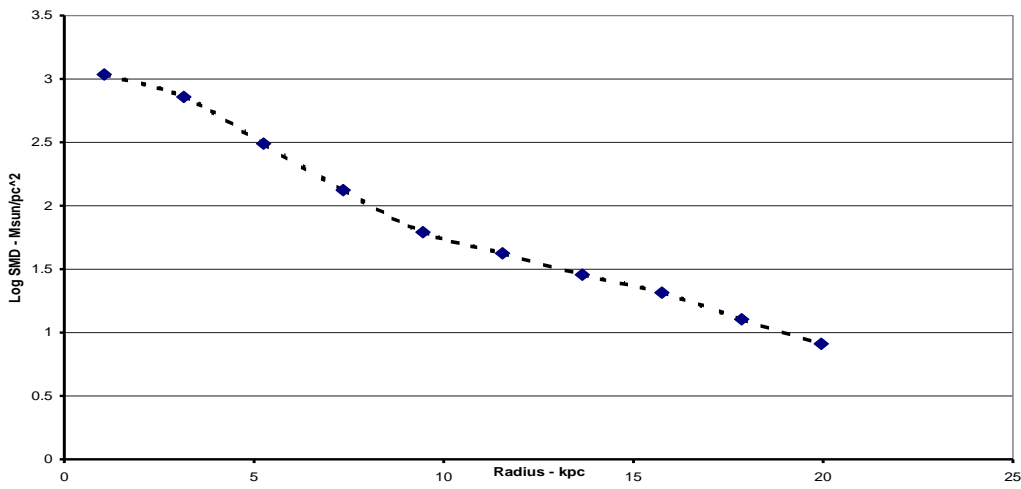


Figure 32c. Surface Mass Density for NGC 3031 Model



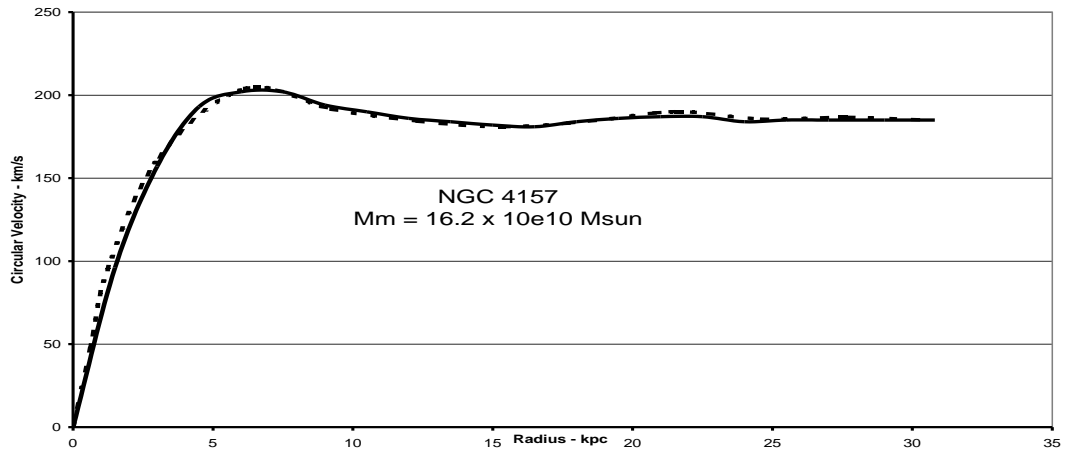


Figure 33a. Velocity Profiles of NGC 4157 and Model

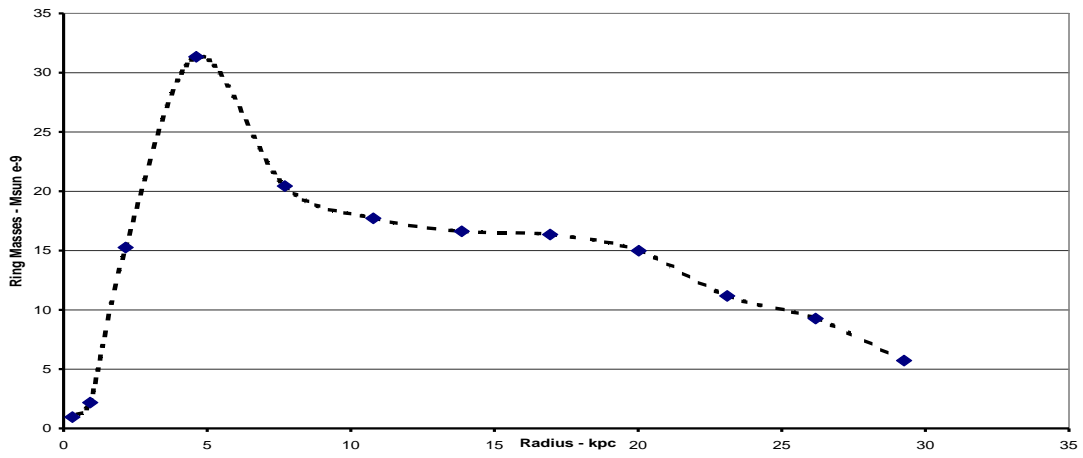


Figure 33b. Ring Mass Distribution for NGC 4157 Model

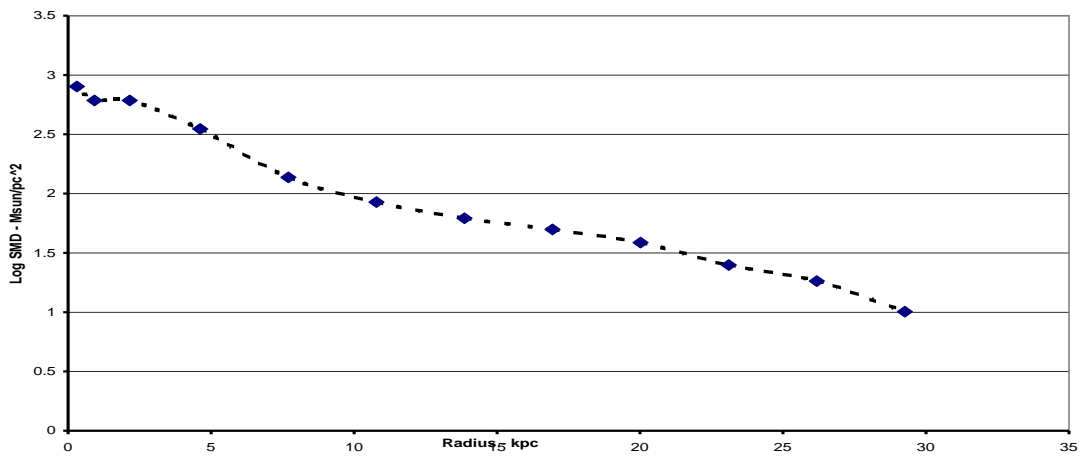


Figure 33c. Surface Mass Density for NGC 4157 Model

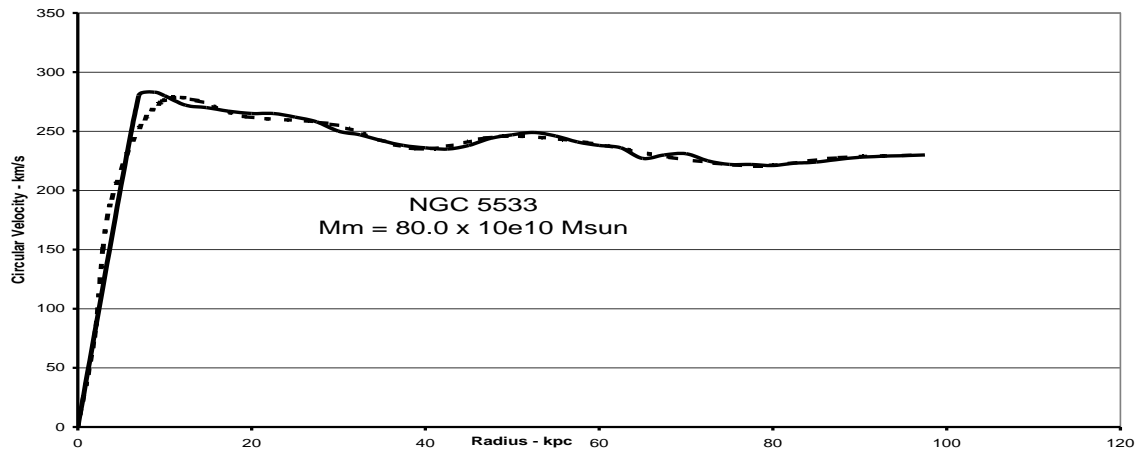


Figure 34a. Velocity Profiles of NGC 5533 and Model

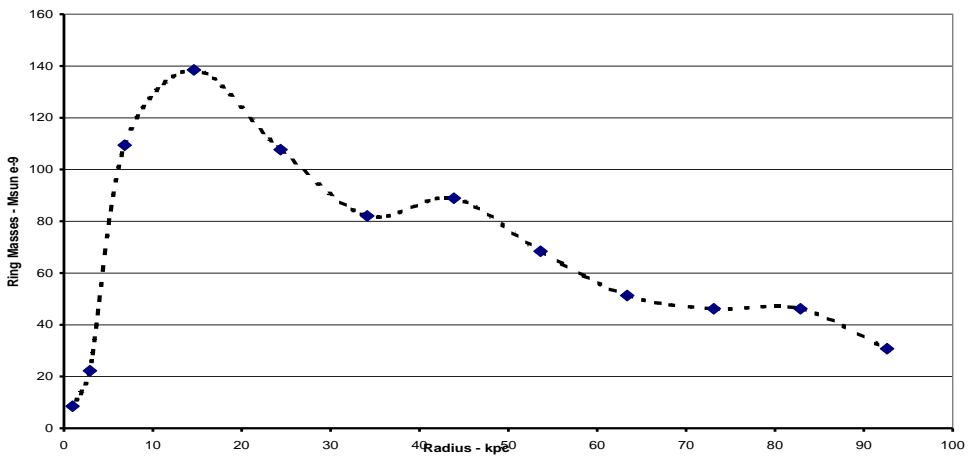


Figure 34b. Ring Mass Distribution for NGC 5533 Model

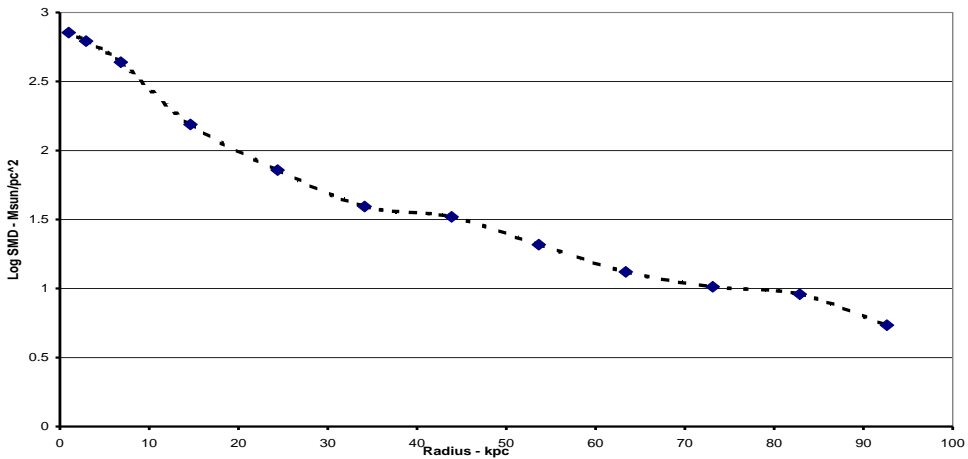
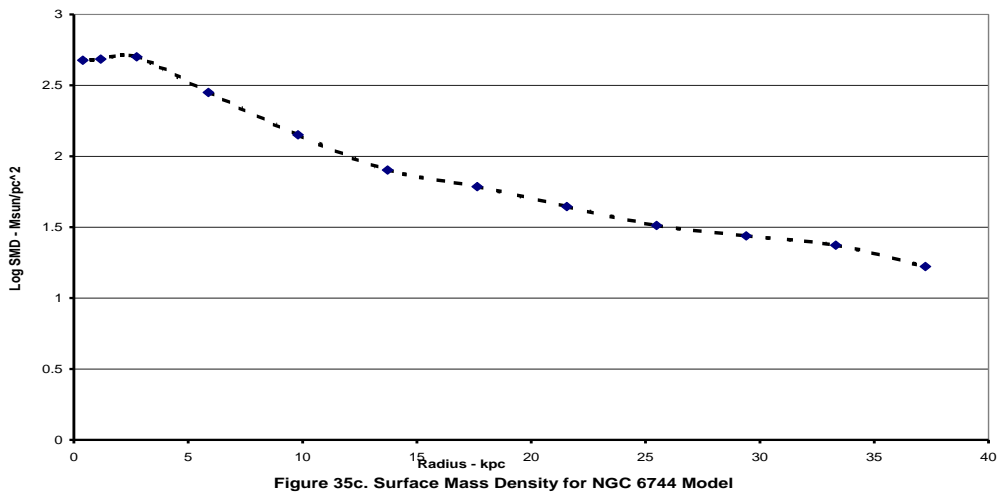
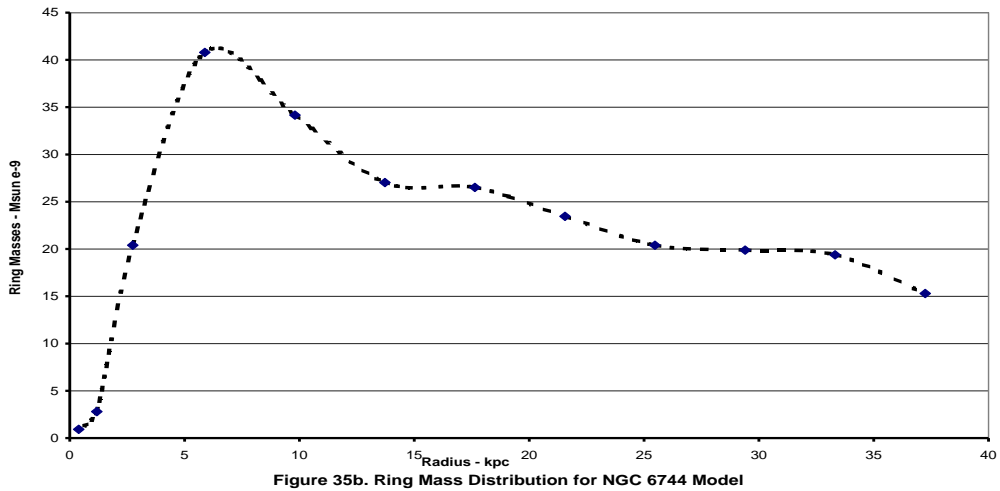
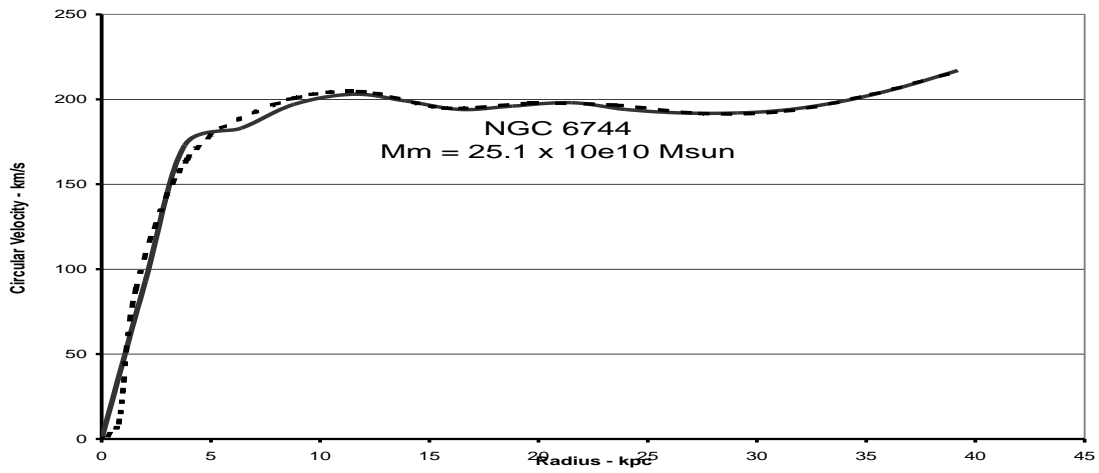
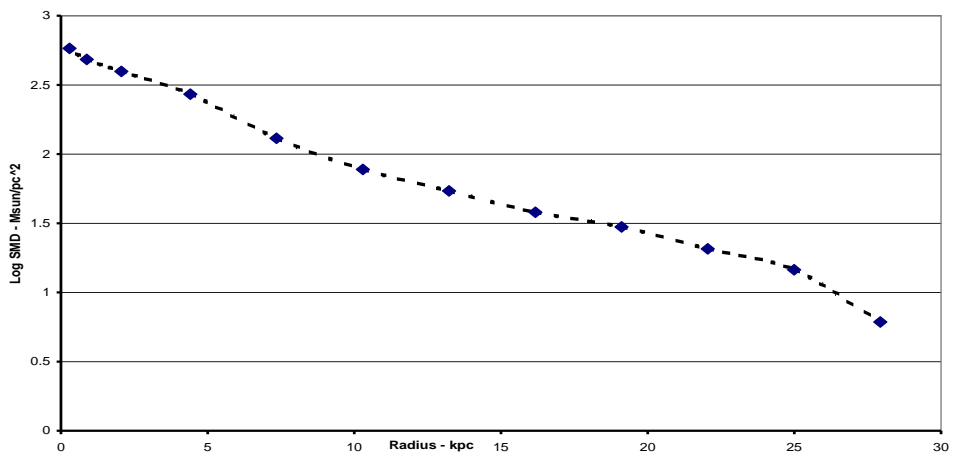
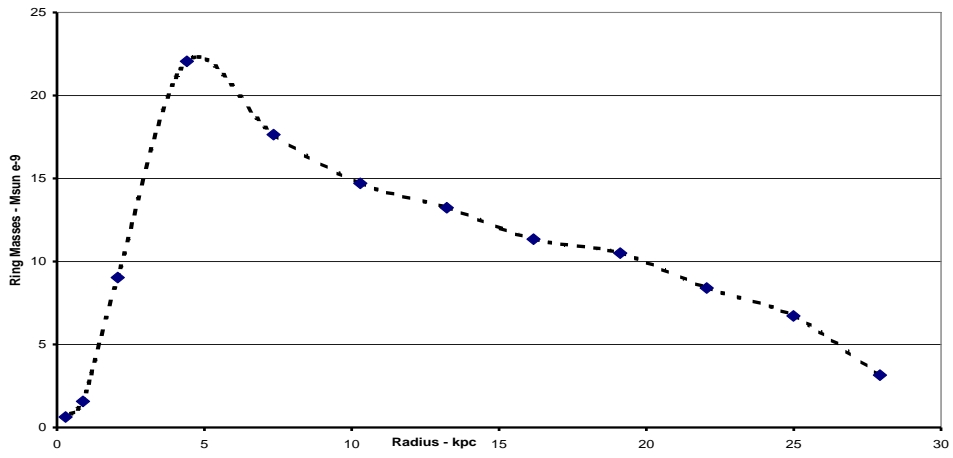
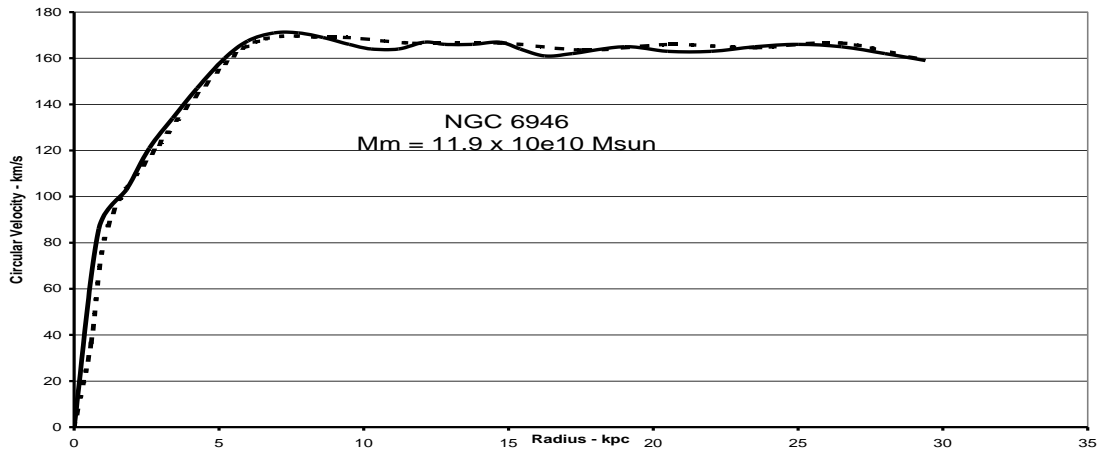


Figure 34c. Surface Mass Density for NGC 5533 Model





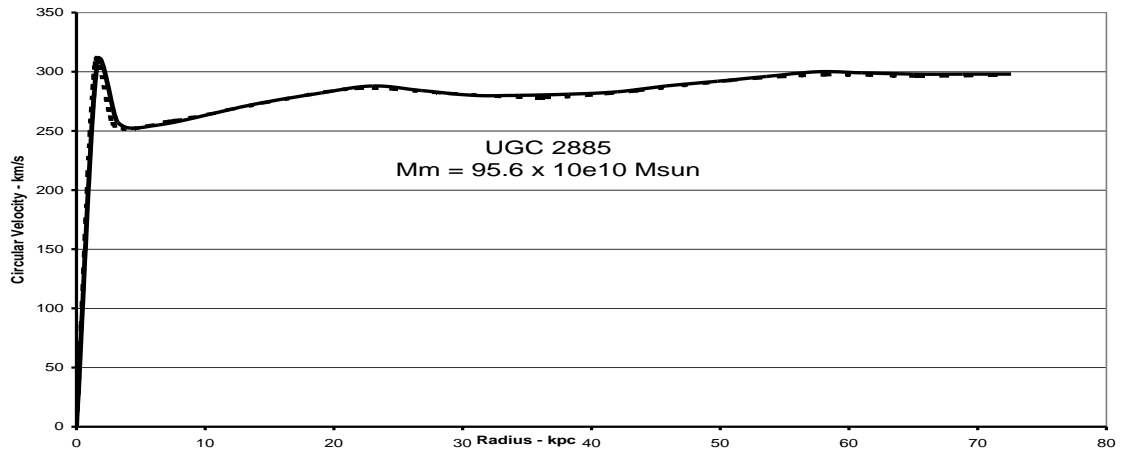


Figure 37a. Velocity Profiles of UGC 2885 and Model

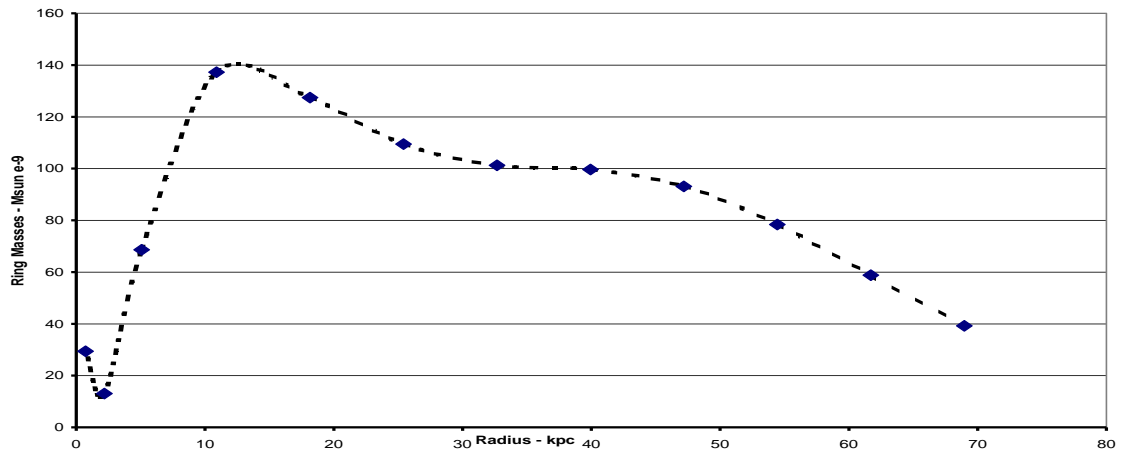


Figure 37b. Ring Mass Distribution for UGC 2885 Model

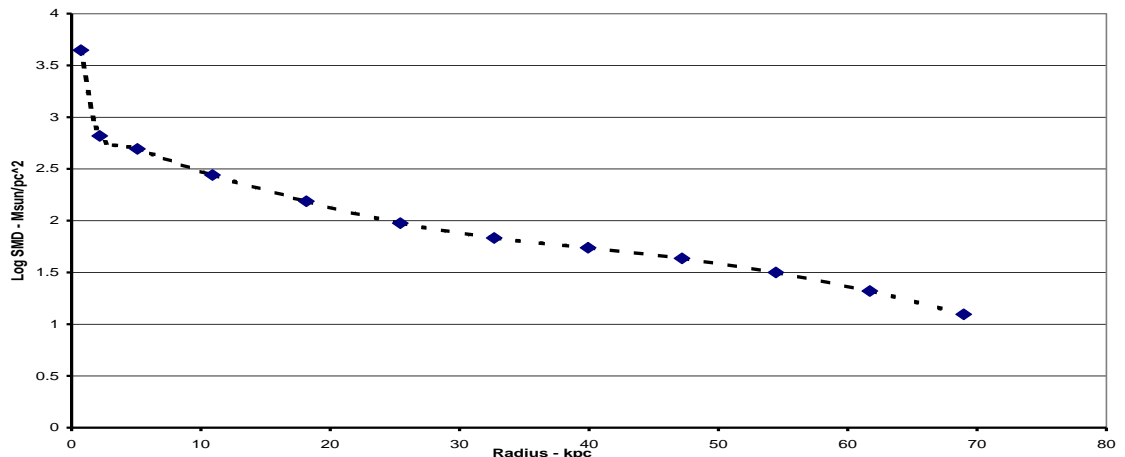


Figure 37c. Surface Mass Density for UGC 2885 Model